



## ABSTRACT

7 Results of four Lagrangian stratocumulus to shallow cumulus transition cases as obtained  
8 from six different large-eddy simulation models are presented. The model output is remark-  
9 ably consistent in terms of the representation of the evolution of the mean state, which is  
10 characterized by a stratocumulus cloud layer that raises with time and which warms and  
11 dries relative to the subcloud layer. Also the effect of the diurnal insolation on cloud-top  
12 entrainment and the moisture flux at the top of the subcloud layer are consistently captured  
13 by the models. For some cases the models diverge in terms of the liquid water path (LWP)  
14 during nighttime which can be explained from the difference in the sign of the buoyancy flux  
15 at cloud base. If the subcloud buoyancy fluxes are positive, turbulence sustains a vertically  
16 well mixed layer causing a cloud layer that is relatively cold and moist and consequently  
17 having a high LWP. After some simulation time all cases exhibit subcloud layer dynamics  
18 which appear to be similar to those of the dry convective boundary layer. The humidity flux  
19 from the subcloud towards the stratocumulus cloud layer, which is one the major sources  
20 of stratocumulus cloud liquid water, is larger during the night than during the day. The  
21 sensible heat flux becomes constant in time whereas the latent heat flux tends to increase  
22 during the transition. These findings are explained from a budget analysis of the subcloud  
23 layer.

## 1. Introduction

Stratocumulus cloud layers are frequently found over relatively cold parts of the subtropical oceans and in the presence of large-scale subsidence. These conditions favor the formation of a thermal inversion, which acts to trap moisture giving rise to extended fields of stratocumulus (Wood 2012). Although the depth of stratocumulus layers is relatively shallow, typically of the order of a few hundreds of meters, they strongly reflect downwelling solar radiation. During the equatorwards transport by the prevailing trade winds over increasing sea surface temperatures the subtropical stratocumulus cloud fields gradually break up and are replaced by shallow cumulus clouds. If a model is not able to capture this stratocumulus to cumulus cloud transition (SCT) this will lead to significant errors in the radiative fluxes received at the ground surface. This is a critical problem as climate models disagree on the change of the subtropical low cloud amount under a global warming scenario, which gives rise to a considerable amount of uncertainty in projections of the future global mean temperature (Bony and Dufresne 2005; Webb et al. 2013; Tsushima et al. 2015).

To investigate the change of the low cloud amount under an idealized warming scenario Zhang et al. (2013) performed experiments with single-column model (SCM) versions of climate models and large-eddy simulation (LES) models. The LES results point to a reduction of the amount of subtropical marine low clouds in a warmer climate (Blossey et al. 2013; Van der Dussen et al. 2015; Bretherton 2015). The study by Zhang et al. (2013), and follow-up studies by Dal Gesso et al. (2014) and Dal Gesso et al. (2015) report a wide scatter in the change of the steady-state subtropical low cloud amount in the SCM results. These results actually give rise to the question how large-scale forcing conditions like the sea surface temperature, free tropospheric temperature and humidity and the large-scale subsidence determine control the SCT.

The SCT has been the subject of several observational (e.g. Albrecht et al. (1995), Bretherton et al. (1995), De Roode and Duynkerke (1997), Sandu et al. (2010)) and mod-

50 elling studies (e.g. Krueger et al. (1995), Sandu and Stevens (2011) and Van der Dussen  
51 et al. (2013)). Chung et al. (2012) studied a series of steady-state LESs in the SCT regime  
52 which can be interpreted as an Eulerian view of the transition. These studies helped to  
53 develop a conceptual view of this transition. According to this model the cloud breakup is  
54 fundamentally driven by the increasing SST. Convective activity driven by surface evapo-  
55 ration increases as the air advects over warmer waters. The strengthening of convectively  
56 driven turbulence enhances the entrainment of warm and dry free-tropospheric air at cloud  
57 top, which leads to a higher virtual potential temperature of the stratocumulus cloud layer  
58 as compared to the subcloud layer. This stratification prevents surface driven thermals to  
59 reach the stratocumulus cloud, except if they become saturated. In that case latent heat  
60 release due to condensation of water allows the plumes to rise as positively buoyant cumulus  
61 clouds, which may penetrate the stratocumulus cloud layer to inject it with moisture from  
62 below (Wang and Lenschow 1995; Miller and Albrecht 1995; De Roode and Duynkerke 1996;  
63 Van der Dussen et al. 2014). Meanwhile, the stratocumulus gradually thins if entrainment  
64 of relatively warm and dry free tropospheric air dominates the longwave radiative cooling  
65 at cloud top and the moisture supply from below. The stratocumulus finally dissipates into  
66 thin and broken patches, penetrated from below by cumulus clouds.

67 To assess whether LES models are capable of faithfully capturing the dynamics of low  
68 clouds, several modeling intercomparison studies have been performed, some of which fo-  
69 cussed on stratocumulus (Moeng et al. 1996; Duynkerke et al. 1999, 2004; Stevens et al. 2005a;  
70 Ackerman et al. 2009), while other studies were dedicated to shallow cumulus (Siebesma et al.  
71 2003; VanZanten et al. 2011), or cumulus penetrating stratocumulus (Stevens et al. 2001).  
72 More recently, four Lagrangian stratocumulus to cumulus transition cases were proposed  
73 to evaluate how well models do in terms of the transition between the two regimes. This  
74 intercomparison study was performed in the framework of the Global Energy and Water  
75 Cycle Exchanges Project (GEWEX) Global Atmospheric System Studies (GASS) and the  
76 European Union Cloud Intercomparison, Process Study & Evaluation Project (EUCLIPSE).

77 Three of the transition cases were based on the "Composite" view of this transition build  
78 using state of the art reanalysis and satellite data (Sandu et al. 2010), while a fourth one  
79 revisited the SCM intercomparison case based on the ASTEX campaign (Bretherton et al.  
80 1999). While ASTEX offers the opportunity to evaluate models against in situ data, the  
81 set of "Composite" transitions represents a more idealized framework for model evaluation,  
82 which offers the possibility of comparing the models for a variety of SCT cases, which differ  
83 for example in terms of amplitude or timescale of the transition.

84 This paper discusses the representation of the four Lagrangian SCT cases in six different  
85 LES models. The Lagrangian approach means that an air mass is followed as it is being  
86 advected by the mean wind from the subtropics towards the equator over an increasingly  
87 warmer SST. Superposed to this change in the surface forcing the air mass is being heated  
88 by absorption of solar radiation during daytime. The paper is organized as follows. In  
89 Section 2 the cases and the LES models are introduced. Section 3 discusses the LES results  
90 with an emphasis on the development of the two-layer structure of the boundary layer.  
91 This decoupled structure motivates to analyse the thermodynamic budgets of the two layers  
92 separately. The contribution of various processes such as entrainment, turbulent fluxes at the  
93 cloud base and radiation to the stratocumulus cloud layer evolution is presented in Section  
94 4. Section 5 analyses the heat and moisture budgets of the subcloud layer and explains the  
95 time evolution of the surface fluxes of heat and moisture. Section 6 discusses and summarizes  
96 the main findings.

## 97 **2. Set-up of the experiments**

98 In this intercomparison case a so-called Lagrangian approach is applied which means  
99 that an air mass is followed as it is being advected by the mean wind allowing to study the  
100 SCT in a single simulation (Schubert et al. 1979). The horizontal advection term in the  
101 conservation equations for heat and moisture may be assumed to be zero in the simulations  
102 as the air parcel is followed along its trajectory. This assumption is acceptable as long as

103 the vertical wind shear is negligibly small.

104 *a. Summary of the Lagrangian stratocumulus transition cases*

105 Three "Composite" cases representing SCTs of varying speed were built based on the  
106 observational study of Sandu et al. (2010). In that study, a large number of Lagrangian  
107 trajectories of air parcels in four subtropical oceans were computed using the wind fields  
108 provided by reanalysis of past observations and the evolution of the cloud and of its environ-  
109 ment along each of these individual trajectories was documented from satellite data sets and  
110 meteorological reanalysis (Moderate Resolution Imaging Spectroradiometer (MODIS), Level  
111 3 data for cloud properties, and European Centre for Medium-Range Weather Forecasts  
112 (ECMWF) Interim Re-Analysis (ERA-Interim, Simmons et al. (2007)) for environmental  
113 properties). This study suggested that averaged forcings can be considered as representative  
114 of individual trajectories, and can therefore be used to initialize numerical simulations of the  
115 transition between the two cloud regimes. Building on these findings, a Composite of the  
116 large-scale conditions encountered along the trajectories for the North East Pacific (NEP)  
117 during June-July-August 2006 and 2007 were used to set up a case study of the SCT, that  
118 will be referred to hereafter as the reference case study and is further described in Sandu and  
119 Stevens (2011). Two variations of this reference case corresponding to a faster, and respec-  
120 tively, to a slower transition in cloud fraction were also derived for the intercomparison study  
121 (and are also described in Sandu and Stevens (2011)). For that, the transitions analyzed for  
122 the NEP during June-July-August 2006 and 2007 were divided into three categories (fast,  
123 intermediate and slow), on the basis of the mean cloud fraction over the first 48 hours. The  
124 initial profiles and the large-scale conditions for each of the three cases represent the medians  
125 of the distributions of the various properties obtained for respective subset of trajectories.

126 The set-up of the fourth SCT case is described in detail by Van der Dussen et al. (2013).  
127 This case is based on observations collected during the first ASTEX Lagrangian experiment  
128 (Albrecht et al. 1995; Bretherton et al. 1995; De Roode and Duynkerke 1997) and large-scale

129 forcing conditions as obtained from ERA-Interim. Since the set-up of the Composite cases  
130 is somewhat idealized, and because the ASTEX case particularly differs from the Composite  
131 cases in terms of precipitation and its relatively cold and moist free troposphere, we think  
132 is useful to discuss its results along with the results from the Composite cases.

133 The initial vertical profiles of the liquid water potential temperature ( $\theta_l$ ), total water spe-  
134 cific humidity ( $q_t$ ) and the horizontal wind velocity components ( $U$  and  $V$ , respectively) for  
135 the four different SCT cases are shown in Fig. 1. The ASTEX case has the smallest value for  
136 the initial inversion jump in the liquid water potential temperature, which gradually increases  
137 in magnitude for the Fast, Reference and Slow cases, respectively. The inversion jumps in  
138 the total specific humidities are also different for each case, with the Slow case having the  
139 driest free atmosphere. The input files provided on the EUCLIPSE website<sup>1</sup> include vertical  
140 profiles of quantities like temperature, humidity and ozone up to the stratosphere, which is  
141 necessary for radiative transfer computations. The transfer of solar radiation is calculated on  
142 the basis of a fixed latitude and longitude. Because the models applied their own radiative  
143 transfer code, the radiative fluxes entering the top of the LES domain differed among the  
144 models, despite that they all used the same prescribed vertical profiles for the atmospheric  
145 column above. The prescribed SST increases with time for each case, which reflects the  
146 Lagrangian equatorwards advection of the simulated air mass (see Fig. 2). The LES models  
147 compute the sensible and latent heat fluxes (SHF and LHF, respectively) from the prescribed  
148 time-dependent SST, a fixed value for the surface roughness length,  $z_0 = 2 \times 10^{-4}$  m, but  
149 each with its own implementation of the Monin-Obukhov similarity theory.

150 For the ASTEX case the large-scale divergence gradually decreases with time, and the  
151 observed weakening of the wind velocities is taken into account by a time-varying geostrophic  
152 forcing (Van der Dussen et al. 2013). For the Composite cases the large-scale divergence and  
153 the geostrophic forcing are constant in time, where the geostrophic winds are the same as  
154 the initial profiles of the horizontal wind velocity components shown in Fig. 1. Although

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<sup>1</sup><http://www.euclipse.nl>

155 the trajectories for the Composite cases are simulated during the same period of time they  
156 have slightly different lengths as their horizontal wind speeds are not the same. The four  
157 Lagrangians also assume a constant surface pressure (see Table 1). The ASTEX and the  
158 three Composite cases last 40 hours and three days, respectively, as these are the timescales  
159 during which the bulk of the transition in cloud cover takes place.

160 *b. Participating large-eddy simulation models and data output*

161 Table 2 lists the models and their acronyms, along with contributors from each partici-  
162 pating group, as well as the main references to the models. The vertical grid resolution in  
163 the lower 540 m is  $\Delta z = 15$  m. To represent the sharp inversion layer capping the cloud  
164 layer the vertical resolution is gradually refined only above this height, and between 645 and  
165 2400 m  $\Delta z = 5$  m. The horizontal domain size is  $4.48 \times 4.48$  km<sup>2</sup> and the number of grid  
166 points in the horizontal directions is  $N_x = N_y = 128$ , implying a horizontal grid spacing  
167  $\Delta x = \Delta y = 35$  m.

168 For each case six large-eddy simulations, each performed with a different code, are pre-  
169 sented. Every code includes a detailed parameterization scheme for radiation and ice-free  
170 cloud microphysical processes, where the latter uses a fixed value for the cloud droplet con-  
171 centration number  $N_d = 100$  cm<sup>-3</sup>.

172 Because the lower tropospheric stability, defined as the difference between the potential  
173 temperature at the 700 hPa pressure level and the ground surface (Klein and Hartmann  
174 1993), is key for the evolution of the SCT, a realistic tendency of the free tropospheric  
175 temperature is needed, in particular as the simulations were performed for a period of two or  
176 three days. Therefore, in contrast to many past studies, all models applied a full radiation  
177 code.

178 To compare the modeling results time series of scalars and hourly-mean vertical profiles  
179 according to the data protocol proposed by VanZanten et al. (2011) were provided by the  
180 modellers. Here it is important to note that liquid water ( $q_l$ ) is defined to include cloud ( $q_c$ )



181 and rain water ( $q_r$ ),  $q_l = q_c + q_r$ , with rain water being defined as drops having a diameter  
182 of  $80 \mu\text{m}$  or larger. In the computation of the cloud fraction and cloud cover a grid cell is  
183 defined to be cloudy if  $q_c > 10^{-5} \text{ kg/kg}$ . Irrespective of whether a model includes rainwater  
184 in its internal representation of the liquid water potential temperature and the total specific  
185 humidity, rain water is included in the profiles of these variables and their fluxes.

### 186 **3. Evolution of the mean state and turbulence structure**

#### 187 *a. Time series*

188 We start our analysis by inspection of the time evolution of the boundary layer, cloud  
189 amount, and the surface fluxes of sensible and latent heat (see Fig. 3). The time variable  
190 in the figure is set such that at the first occasion of local noon  $t = 0$ . Nighttime periods  
191 (denoted by 'N1', 'N2' and 'N3' at the top of Fig. 3h) are indicated by the grey vertical  
192 bands in the plots according to the simulation periods summarized in Table 3. For each  
193 LES model, and for each daytime and nighttime period we calculated time-mean results. To  
194 get an appreciation of the spread in the modelling results, Table 4 presents the overall LES  
195 mean values and standard deviations. Note that because during the first two hours of the  
196 simulations the turbulence has not fully developed yet, the results during this spin-up period  
197 were not used.

198 In brief, the results show that for all cases the cloud-topped boundary layer is gradually  
199 deepening with time, while the cumulus cloud base height reaches an approximate steady  
200 state. The effect of the diurnal variation of the solar radiation is clearly found from the time  
201 series of the LWP. Due to the absorption of solar radiation in the cloud layer the LWP has  
202 reduced values during daytime. The cloud layer breaks up during the second daytime period  
203 'D2' for the Fast case, although it tends to recover to a closed cloud deck during the second  
204 night 'N2', except for MOLEM. The Slow case appears to maintain an almost closed cloud  
205 deck during the entire simulation period. For all SCTs the entrainment velocity is much

206 larger during nighttime than during the day. Finally, for the Composite cases the surface  
207 evaporation gradually increases whereas the sensible heat flux remains rather small.

208 A closer inspection reveals that during local noon the growth of the inversion height  
209 becomes very small for the Composite cases, which is due to a reduced cloud-top entrainment  
210 rate whereas the subsidence keeps pushing down the boundary-layer top (see Figs. 3a-d).  
211 The variation in the boundary layer depth as represented by the standard deviation  $\sigma_{z_i}$   
212 computed from the six model results also gradually increases with time (see Table 4). Given  
213 the myriad of physical processes that control the boundary layer depth (e.g. turbulence,  
214 radiation, entrainment and drizzle), the values of  $\sigma_{z_i}$  can be considered as relatively small,  
215 with maximum values of 100 m except for the Fast case which gives a value of 200 m during  
216 the third nighttime period 'N3'. The height of the lowest cumulus cloud base ( $z_{\text{cu,base}}$ ) is  
217 very consistently represented among the models, its standard deviation being less than 50 m.  
218 We find an overall relatively small increase of  $z_{\text{cu,base}}$  during the first part of the simulations,  
219 and during the second part it becomes almost constant in time.

220 By contrast, the intermodel spread in the cloud liquid water path (LWP) is relatively large  
221 particularly during the night (see Figs. 3e-h) similar to what was found in the stratocumulus  
222 model intercomparison study by Stevens et al. (2005a). The LES agree fairly well in terms  
223 of the representation of the diurnal variation of the LWP, although the amplitude is larger  
224 in the MPI/UCLA, DHARMA and EULAG models. The latter model explains a significant  
225 part of  $\sigma_{\text{LWP}}$ , which is relatively large as compared to the mean value, in particular during  
226 nighttime.

227 MOLEM and EULAG have a consistently different longwave radiative forcing for the  
228 three Composite cases as compared to the other LES models which results are very similar.  
229 For example, during the first night of the Composite cases the longwave radiative flux di-  
230 vergence in the cloud layer is about  $5 \text{ Wm}^{-2}$  smaller in MOLEM and about  $10 \text{ Wm}^{-2}$  larger  
231 in EULAG. The effect of the differences in the longwave radiative cooling on the cloud layer  
232 evolution is discussed in detail in Section 5. Figs. 3i-l show the time evolution of the cloud

233 cover. Only in the EULAG model a solid cloud is maintained for all SCTs which possibly  
234 results from the imposed stronger cloud longwave radiative cooling. In the other models the  
235 stratocumulus starts to break up some hours after sun rise due to the absorption of solar  
236 radiation in the cloud layer (Nicholls 1984). Most of the time the stratocumulus is able to  
237 recover to a closed-cell cloud deck after sunset. The difference between the three Compos-  
238 ite cases becomes clear as the cloud cover tends to reduce more rapidly for the Fast case  
239 compared to the Reference or Slow cases, which is in a rough agreement with estimations  
240 of cloud cover from MODIS. However the inter-model differences in the daytime cloud cover  
241 are rather large. For example, for the Fast and Reference cases the standard deviation of  
242 the cloud cover has maximum values during the third daytime period 'D3'.

243 The absorption of the solar radiation leads to the warming and the thinning of the cloud  
244 layer. The absorption of solar radiation in the cloud layer counteracts the longwave radiative  
245 cooling at the cloud top. The stabilization of the cloud layer during daytime tends to weaken  
246 the buoyancy production of turbulence, which in turn causes a reduction in the entrainment  
247 velocity. If we compare the entrainment velocity for the four cases, we find smaller values  
248 for a stronger thermal stratification as measured by the inversion jump values of  $\theta_1$ . During  
249 the first nighttime period 'N1' the entrainment rate is largest for the ASTEX case, and  
250 gradually becomes smaller for the Fast, Reference and Slow cases, respectively. There is a  
251 good agreement in the modelled entrainment velocity, with a maximum standard deviation  
252 of about  $1 \text{ mm s}^{-1}$  (see Figs. 3m-p and Table 4).

253 The LES models give SHF values that are less than  $10 \text{ Wm}^{-2}$  (see Figs. 3q-t). The  
254 LHF tends to increase with time (see Figs. 3u-x), except for the ASTEX case for which  
255 a flattening of the temporal SST increase and a weakening geostrophic forcing yields lower  
256 wind velocities and consequently lower LHF values. The Composite cases exhibit a gradual  
257 increasing trend in the LHF, with an imposed diurnal cycle in which the flux increases faster  
258 during the night than during the day. The standard deviation of the LHF is within 10  
259  $\text{Wm}^{-2}$ .

260 Although the bulk features of the time variation of the cloud structure and the differences  
261 between the four cases are consistently represented, the variation in the cloud cover and the  
262 LWP leads to a rather large value for the standard deviation of the net shortwave radiation  
263 at the surface, with a maximum value of  $80 \text{ Wm}^{-2}$  during the third daytime period 'D3'  
264 for the Fast case. During the entire simulation period the standard deviation of the net  
265 longwave radiation at the surface is within  $10 \text{ Wm}^{-2}$ .

266 *b. Boundary-layer decoupling*

267 Hourly mean vertical profiles of  $\theta_1$  and  $q_t$  obtained from the Fast case 48 hours from  
268 local noon are shown in Fig. 4. The stratocumulus layer has a higher  $\theta_1$  and a lower  $q_t$   
269 than the subcloud layer. The subcloud and cloud layer each are rather well mixed vertically.  
270 The lowest inversion height is found in MOLEM, and the stronger longwave radiative cloud  
271 layer cooling imposed in the EULAG model causes a much higher inversion layer height  
272 due to a larger entrainment rate (Figs. 3m-p). At this time, all models show a broken  
273 stratocumulus cloud deck, with the cloud fraction varying roughly between 0.05 and 0.78,  
274 except for EULAG which maintains an almost closed cloud deck for all the SCTs. The  
275 differences in the horizontal wind velocity components across the inversion are small. This  
276 is also the case for the Slow and Reference cases where the jumps are smaller than  $2 \text{ ms}^{-1}$ .

277 The different evolutions in  $\theta_1$  and  $q_t$  in subcloud and cloud layers are illustrated in Fig.  
278 5. We use the subscript 'ml' to denote the subcloud mixed-layer mean value. It is computed  
279 from the mean between the first level above the surface and the cumulus cloud base height  
280  $h$ . Likewise we use the subscript 'cld' to indicate the stratocumulus mean value between its  
281 mean base and top heights. As an easy reference the values at the surface and just above the  
282 inversion are also shown in the figure, and are indicated by subscripts 'sfc' and  $z_i^+$ , respec-  
283 tively. The mean values of  $\theta_1$  in the subcloud and stratocumulus cloud layers both increase  
284 in time, with  $\theta_{1,\text{ml}}$  roughly following the trend of the surface value, and  $\theta_{1,\text{cld}}$  increasing at  
285 a slightly faster pace. In contrast to  $q_{t,\text{ml}}$ ,  $q_{t,\text{cld}}$  shows a drying trend, which implies that

286 the drying of the stratocumulus cloud layer by entrainment and drizzle is stronger than the  
 287 moisture input by the updrafts from the subcloud layer.

288 After some simulation time the vertical profiles of  $\theta_1$  and  $q_t$  all resemble a decoupled  
 289 boundary layer structure, with a cloud layer that is relatively warm and dry with respect  
 290 to the subcloud layer (Nicholls 1984; Bretherton and Wyant 1997; Stevens 2000; Wood and  
 291 Bretherton 2004). A convenient way to measure the degree of decoupling is given by Park  
 292 et al. (2004) who defined the following decoupling parameter,

$$\alpha_\psi = \frac{\psi_{\text{cld}} - \psi_{\text{ml}}}{\psi(z_i^+) - \psi_{\text{ml}}}, \quad (1)$$

293 with  $\psi \in \{\theta_1, q_t\}$ , and  $z_i^+$  the height just above the inversion layer. The decoupling parameter  
 294 is equal to zero if the boundary layer is well mixed, i.e.  $\theta_1$  and  $q_t$  constant with height.

295 Fig. 6 compares the decoupling parameters  $\alpha_{q_t}$  and  $\alpha_{\theta_1}$  as found from the LES results with  
 296 a fit of  $\alpha_{q_t}$  that was obtained from aircraft observations analyzed by Wood and Bretherton  
 297 (2004), their Fig. 5. Both the observations and the LES results suggest a stronger decoupling  
 298 for deeper boundary layers, as measured by larger values of  $\alpha_{q_t}$  and  $\alpha_{\theta_1}$ . The results presented  
 299 in Table 2 of Wood and Bretherton (2004) appear to give a somewhat smaller difference  
 300 between  $\alpha_{q_t}$  and  $\alpha_{\theta_1}$  than the LES results.

301 Large values of the decoupling parameters indicate that the cloud layer is relatively warm  
 302 and dry with respect to the subcloud layer. Because a high temperature or a low total water  
 303 amount in the cloud tend to reduce the cloud liquid water content, we will now take a closer  
 304 look at the time evolution of the decoupling parameters. In particular, we will inspect the  
 305 results for the Slow case which shows a rather large scatter in the nighttime LWP values  
 306 among the six LES. The gradual deepening of the boundary layer is reflected in the gradual  
 307 increase of  $\alpha_{q_t}$  and  $\alpha_{\theta_1}$  with time (see Fig. 7). However, during the first nighttime period the  
 308 boundary layer gets back to a very well mixed vertical structure, while during the second  
 309 nighttime period a strong variation in the degree in the decoupling is observed. For the latter  
 310 period the DHARMA and MPI/UCLA models show an almost perfectly vertically mixed  
 311 boundary layer, whereas the boundary layer remains rather strongly decoupled in DALES.

312 Inspection of the LWP values confirm its strong correlation with the degree of decoupling,  
313 with DHARMA and MPI/UCLA having the largest LWP values and the smallest values for  
314 the decoupling parameters.

315 *c. Turbulence*

316 The  $\theta_1$  and  $q_t$  fields presented in Fig. 8 show a distinct three layer structure with a very  
317 sharp inversion layer that separates the stratocumulus layer from the dry free troposphere.  
318 The top of the subcloud layer itself is much more diffuse. The encircled numbers '1' and '2'  
319 are near rising subcloud plumes that become saturated and ascend further as cumulus clouds  
320 thereby transporting subcloud layer moisture towards the stratocumulus. Interestingly, area  
321 '3' is in an area above cumulus clouds, and shows sinking motions near two holes in the  
322 stratocumulus cloud deck, that resulted from evaporation of cloud water by entrainment of  
323 free tropospheric air (Gerber et al. 2005; De Roode and Wang 2007; de Lozar and Mellado  
324 2015). Turbulence in clear air patches above the subcloud layer was also detected from  
325 aircraft observations during ASTEX (De Roode and Duijnkerke 1996).

326 The findings presented so far suggest that the inter-model spread in the LWP during  
327 nighttime can be linked to the various strengths of the decoupling between the cloud and  
328 the subcloud layer. Stevens et al. (2005b) reported similar findings for the DYCOMS-II  
329 nighttime stratocumulus LES intercomparison case. They found a strong link between the  
330 buoyancy flux profile, the vertical velocity variance, and the degree of decoupling. It is  
331 therefore instructive to repeat their analysis by inspecting the turbulence profiles for the  
332 SCTs. Fig. 9 shows hourly-mean vertical profiles of the vertical velocity variance  $\overline{w'w'}$ , the  
333 virtual potential temperature flux  $\overline{w'\theta'_v}$ , the vertical flux of total water specific humidity  $\overline{w'q'_t}$ ,  
334 and the turbulent kinetic energy (TKE) for the Slow case. Note that the fluxes of the virtual  
335 potential temperature and the buoyancy  $b$  are proportional,  $\overline{w'b'} = \beta\overline{w'\theta'_v}$ , with  $\beta = g/\theta_0$ ,  $g$   
336 the acceleration due to gravity and  $\theta_0$  a constant reference temperature.

337 The surface buoyancy fluxes are positive. Towards the top of the subcloud layer the

338 buoyancy flux decreases and can even become negative, indicating that, on average, rising  
 339 plumes are negatively buoyant. If the plumes become saturated with water vapor, the latent  
 340 heat release due to condensation enables them to rise further as positively buoyant clouds.  
 341 The negative buoyancy fluxes just above the top of the cloud layer are due to entrainment  
 342 of warm free tropospheric air. Longwave radiative cooling in the cloud top regions leads to  
 343 buoyancy production, and as the cooled cloud parcels become heavier than the surrounding  
 344 air they start sinking, leading to a positive buoyancy flux.

345 The imposed solar radiative heating of the cloud layer during daytime has a distinct  
 346 effect on the turbulence structure of the boundary layer. In particular, the signature of  
 347 a decoupled boundary layer structure is clearly visible from the double peak structure in  
 348  $\overline{w'w'}$  and the rather low values for the TKE. As was observed at the end of the ASTEX  
 349 Lagrangian (De Roode and Duynkerke 1996), the vertical profiles for the buoyancy flux and  
 350 the vertical velocity variance during daytime and in the final stages of the Composite SCTs  
 351 become similar to ones found in the dry convective boundary layer (Stevens 2007). Although  
 352 this decoupled two-layer turbulence structure might be considered as a difficult condition  
 353 to be represented by the LES models, there is a much better agreement in the turbulence  
 354 profiles during daytime than during the night. For example, the differences in terms of  $\overline{w'w'}$   
 355 profiles and TKE is much larger during nighttime. At first sight this seems at odds with  
 356 the nighttime buoyancy fluxes which appear to agree pretty well. If we however zoom in  
 357 at models that have slightly positive buoyancy fluxes at the top of the subcloud layer, for  
 358 example MPI/UCLA and DHARMA, we find that they have the largest  $\overline{w'w'}$  and TKE  
 359 values. Stated more precisely, at 36 hrs from local noon their  $\overline{w'w'}$  profiles have a single  
 360 peak in contrast to the other models that tend towards a double peak structure.

361 Bretherton and Wyant (1997) argued that the buoyancy flux at the top of the subcloud  
 362 layer  $\overline{w'\theta'_{v,h}}$  is key to the development of a decoupled boundary layer. Because the sign of  
 363  $\overline{w'\theta'_v}$  determines whether turbulence will be diminished or amplified, Figs. 10a-d present the  
 364 time evolution of the flux ratio  $r_{\theta_v}$ , which defines the flux at the top of the subcloud layer  $h$

365 normalized by its surface value,

$$r_{\theta_v} = \frac{\overline{w'\theta'_{v,h}}}{\overline{w'\theta'_{v,sfc}}}. \quad (2)$$

366 Table 5 shows the mean values of  $r_{\theta_v}$  for the daytime and nighttime periods. In particular  
 367 during nighttime periods with positive  $r_{\theta_v}$  values the boundary layer is found to be vertically  
 368 well-mixed, whereas a negative  $r_{\theta_v}$  is indicative for decoupling as characterized by a double  
 369 peak structure in the vertical velocity variance profile. For the Composite cases all models  
 370 quickly obtain a small or negative  $r_{\theta_v}$  for the Fast case, whereas for the Slow case two models  
 371 return to a positive  $r_{\theta_v}$  during the second nighttime period 'N2'. However, similar to the  
 372 daytime periods, at the end of the simulations the boundary layer becomes permanently  
 373 decoupled, as indicated by  $r_{\theta_v}$  which remains negative during the third nighttime period  
 374 'N3', except for EULAG.

375 Likewise the flux ratio  $r_{q_t}$  is defined similar to  $r_{\theta_v}$ , and measures the fraction of the  
 376 surface evaporation which is transported out of the subcloud layer. Figs. 10e-h and the  
 377 conditionally sampled results in Table 5 show that  $r_{q_t}$  exhibits a clear diurnal cycle. During  
 378 daytime  $r_{q_t} < 1$ , which indicates that moisture accumulates in the subcloud layer, whereas  
 379 during the first nighttime period the rate at which cumulus clouds transport water out of  
 380 the subcloud layer exceeds the surface evaporation leaving a drying of the subcloud layer.  
 381 In general  $r_{q_t}$  is larger during nighttime as compared during daytime.

## 382 4. Stratocumulus LWP budget

383 To understand what controls the LWP evolution and what leads to the LWP differences  
 384 among the LES models we have assessed the effect of turbulence, radiation and drizzle on  
 385 the LWP evolution, following its budget analysis by Van der Dussen et al. (2014),

$$\frac{\partial LWP}{\partial t} = \text{Ent} + \text{Base} + \text{Rad} + \text{Prec} + \text{Subs}. \quad (3)$$

386 As noted by Ghonima et al. (2015) this budget equation is analogous to the cloud layer depth  
 387 budget by Wood (2007) and is derived from the conservation equations for heat, water and



388 mass, and the terms are defined by

$$\begin{aligned}
\text{Ent} &= \rho w_e (\eta \Delta q_t - \Pi \gamma \eta \Delta \theta_1 - h_{\text{cld}} \Gamma_{q_1}) \\
\text{Base} &= \rho \eta [\overline{w' q'_t}(z_b) - \Pi \gamma \overline{w' \theta'_1}(z_b)] \\
\text{Rad} &= \frac{\eta \gamma}{c_p} [F_{\text{rad}}(z_t) - F_{\text{rad}}(z_b)] \\
\text{Prec} &= -\rho [P(z_t) - P(z_b)] \\
\text{Subs} &= -\rho h_{\text{cld}} \Gamma_{q_1} \overline{w}(z_t),
\end{aligned} \tag{4}$$

389 with  $\rho$  the density of air,  $\eta$  and  $\gamma$  are factors that include the Clausius-Clapeyron relation,  
390  $c_p$  is the specific heat for dry air,  $\Gamma_{q_1} < 0$  is the lapse rate of the liquid water specific  
391 humidity,  $\Pi$  is the Exner function,  $P$  the drizzle rate, and the stratocumulus cloud layer  
392 depth  $h_{\text{cld}} = z_t - z_b$ , where the heights of the mean stratocumulus cloud base  $z_b$  and cloud  
393 top  $z_t$  were diagnosed from the heights between which the cloud fraction is larger than 0.4.  
394 The thermodynamic factors arise because if the cloud layer is moistened, the release of heat  
395 due to condensation of water causes the temperature to rise, which enhances the saturation  
396 specific humidity, such that not the full amount of the added moisture becomes liquid. A  
397 similar arguments holds if heat is added to the cloud layer, as its warming effect will act to  
398 evaporate some liquid water causing a compensating cooling effect.

399 The turbulent flux at the top of the cloud layer has been substituted by the flux-jump  
400 relation (Lilly 1968), which states that the flux of a quantity  $\psi$  at the top of the boundary  
401 layer is proportional to the entrainment velocity and the jump of the quantity across an  
402 infinitesimally thin inversion layer, e.g. for  $q_t$ ,

$$\overline{w' q'_t}_{z_t} = -w_e \Delta \overline{q}_t. \tag{5}$$

403 Application of this relation gives a more accurate estimation of the flux of  $\theta_1$  at the top of the  
404 cloud layer, as the diagnosed slab-averaged Reynolds-averaged flux typically underestimates  
405 the entrainment flux due to the fact that the inversion layer has a finite depth. The inversion  
406 jumps of  $\theta_1$  and  $q_t$  are shown in Fig. 11. The LWP budget analysis for the ASTEX case has

407 been reported by Van der Dussen et al. (2016) to investigate why a reduction of the large-  
 408 scale subsidence causes the stratocumulus cloud deck to persist longer despite an increase  
 409 in the entrainment velocity. This study also demonstrated that a very good correspondence  
 410 can be obtained between the actual and the LWP tendency as diagnosed from the rhs  
 411 of (4). Because our analysis is based on hourly-mean processed data of fluxes and mean  
 412 quantities, the residual in the LWP budget is larger than in Van der Dussen et al. (2016).  
 413 Nevertheless some robust features emerge from the dominant LWP budget terms shown in  
 414 Fig. 12 and the corresponding Table 6, which shows the mean results and the standard  
 415 deviations during a full daytime or nighttime period. We note that model results were not  
 416 used if stratocumulus was not detected during some part of the selection period. Specifically,  
 417 for the Fast case stratocumulus disappeared in MOLEM and DHARMA during periods 'D2'  
 418 and 'D3', respectively, and did not recover, while DALES and MPI/UCLA temporarily had  
 419 no stratocumulus during period 'D3'. For the Reference case MOLEM had no stratocumulus  
 420 during 'D3'. For ASTEX no stratocumulus was present for DALES, SAM, MOLEM and  
 421 DHARMA during period 'D2'.

422 The entrainment drying and warming effects are represented by  $\text{Ent}_{\text{dry}}$  and  $\text{Ent}_{\text{heat}}$  (the  
 423 first two terms on the rhs of Ent), and likewise the Rad term has been split in a longwave  
 424 and a shortwave contribution,  $\text{Rad}_{\text{LW}}$  and  $\text{Rad}_{\text{SW}}$ , respectively. Longwave radiative cooling  
 425 and cloud base moisture fluxes are the dominant terms which support the increase of the  
 426 LWP. During daytime absorption of solar radiation tends to diminish the LWP. Its cloud  
 427 layer warming effect acts to stabilize the cloud layer with respect to the subcloud layer, and  
 428 as a result the input of moisture from below the cloud layer diminishes. Secondly, as the  
 429 solar warming counteracts the destabilization due to longwave cooling at the cloud top, the  
 430 cloud layer thinning due to entrainment of relatively warm and dry air also decreases. If the  
 431 cloud layer becomes sufficiently thin or broken, we find that the longwave radiative cooling  
 432 also strongly decreases. The EULAG model has the strongest longwave cooling effect, which  
 433 apparently prohibits the stratocumulus cloud layer to break up for the Composite cases

434 (see Fig. 3). Note that the state of the atmospheric column above the LES domain was  
 435 prescribed for all cases, and the differences in the downward radiative fluxes at the top of  
 436 the LES domain are therefore due to different radiative transfer schemes used in the LES  
 437 models.

438 The budget analysis indicates that the imbalance of a couple of rather large contributions  
 439 to the LWP tendency determines the actual LWP tendency. It also clarifies the role of en-  
 440 trainment. The Fast case has the smallest inversion jumps of  $\theta_1$  as compared to the Reference  
 441 and Slow cases. Because of this relatively weak thermal stability it has the largest entrain-  
 442 ment rates, resulting in the largest cloud thinning effects due to the mixing of relatively  
 443 warm and dry air from just above the inversion.

444 The cloud thinning effect due to precipitation is very small except for the ASTEX case  
 445 during the first nighttime period. The difference in drizzle between the ASTEX and the  
 446 Composite cases can be understood qualitatively from a drizzle parameterization at the  
 447 cloud base height derived from observations by Comstock et al. (2004),

$$P_{\text{cb}} = 0.37 \left( \frac{LWP}{N_d} \right)^{1.75}, \quad (6)$$

448 which thus depends on the LWP and the cloud droplet concentration number  $N_d$  which is set  
 449 to  $100 \text{ cm}^{-3}$  in the simulations. VanZanten et al. (2005) derived a similar relation. The three  
 450 Composite cases have typical maximum LWP values of the order of  $100 \text{ gm}^{-2}$ , for which the  
 451 parameterization above gives a drizzle rate of  $11 \text{ Wm}^{-2}$ . For higher LWP values such as  
 452 found for the ASTEX case, the drizzle rate becomes more significant too, with values of 38  
 453 and  $77 \text{ Wm}^{-2}$  for LWP values of 200 and  $300 \text{ gm}^{-2}$ , respectively. The ASTEX case is the  
 454 only simulation which starts during nighttime during which the stratocumulus cloud tends  
 455 to thicken. It also has a rather cold and moist free troposphere, which tends to weaken its  
 456 capability to thin the stratocumulus layer by entrainment.

457 Van der Dussen et al. (2013) showed from additional sensitivity experiments for the  
 458 ASTEX case that the difference in the LWP is mainly attributable to differences in the  
 459 precipitation rate. They also found that stronger precipitating stratocumulus had less en-

460 trainment of warm and dry inversion air at its top. During daytime model differences in  
 461 LWP are also diminished by solar radiative heating of the cloud layer. This mechanism  
 462 is particularly clear during the third daytime period 'D3' of the Fast and Reference case  
 463 simulations by EULAG. The LWP in this model is much higher than in the others (see Figs.  
 464 3f and g), which causes a much stronger cloud thinning tendency due to the absorption of  
 465 solar radiation (Figs. 12f and g).

## 466 5. Subcloud layer heat and moisture budgets

467 The behaviour of the surface SHF and LHF during the transitions is very different in  
 468 the sense that the SHF becomes approximately constant at about  $10 \text{ Wm}^{-2}$ , whereas the  
 469 LHF tends to increase with time during the Lagrangian advection of the cloudy air mass  
 470 (see Figs. 3q-x). A classical framework to explain the time evolution of surface fluxes is  
 471 the mixed-layer model (MLM), which assumes a vertically well-mixed boundary layer (Lilly  
 472 1968; Schubert et al. 1979; Nicholls 1984). The values of the decoupling parameters  $\alpha_{q_t}$  and  
 473  $\alpha_{\theta_1}$  indicate that this assumption is not appropriate for relatively deep boundary layers. On  
 474 the other hand, since the subcloud layer is vertically well mixed, the MLM framework may  
 475 be applied to this lower part of the boundary layer.

### 476 a. Evolution of the subcloud-layer height

477 Figs. 3a-d show that the gradual increase of the subcloud layer height, which approxi-  
 478 mately coincides with the cumulus cloud base height, reduces significantly during the final  
 479 stages of the simulations. The time evolution of the subcloud mixed layer height  $h$  can be  
 480 expressed in terms of the mass budget equation (Neggers et al. 2006),

$$\frac{\partial h}{\partial t} = E + \bar{w}|_h - M, \quad (7)$$

481 where  $E$  is a positive term that represents the entrainment process which mixes air into the  
 482 subcloud layer from above,  $\bar{w}|_h$  is the large-scale vertical velocity at the top of the subcloud

483 layer which is negative for the cases considered here, and  $M > 0$  is related to the shallow  
484 cumulus mass flux which acts as a sink term. Because the relative humidity (RH) in a  
485 vertically well mixed layer increases with height, an initial deepening of the subcloud layer  
486 depth  $h$  will subsequently lead to higher RH values at its top. This will trigger shallow  
487 cumulus clouds whose mass flux will reduce the height of the mixed layer, and hence the RH  
488 at its top. In this way the cumuli act as a kind of valve that will maintain an approximate  
489 constant RH at the top of the subcloud layer (Bretherton et al. 2004).

490 *b. Analysis of the results*

491 To study the behaviour of the surface heat fluxes we will apply a mixed layer model to  
492 the subcloud layer. This model assumes a quasi-steady state, which means that temporal  
493 changes in conserved thermodynamic variables are constant with height. This allows to  
494 obtain simple solutions for the vertical fluxes, which in this framework only depend on the  
495 values at the bottom and the top of the mixed layer, and the net effect of diabatic processes.  
496 In fact, if we approximate the mixed-layer height to be constant in time, and if we express  
497 the sea surface temperature as a linear function of time, it is possible to obtain analytical  
498 expressions for the thermodynamic evolution of the subcloud layer provided that we close  
499 the system with use of the flux ratios  $r_{\theta_v}$  and  $r_{q_t}$ , respectively. Table 7 presents the notation  
500 for the initial conditions, the time-dependent surface boundary conditions, the definitions of  
501 the time scales of the system as derived in the Appendix, and the constants  $C_1$ ,  $C_2$  and  $C_3$ .  
502 In particular, we find that the mixed-layer values for  $\theta_v$  and  $q_t$  change in time according to,

$$\theta_{v,ml}(t) = \gamma_{\theta_v} t + C_1 + C_2 \exp^{-t/\tau_{\theta_v}}, \quad (8)$$

503

$$q_{t,ml}(t) = \frac{q_{sat,sfc,0}}{1 + \frac{\tau_q}{\tau_{CC}}} \exp^{t/\tau_{CC}} + C_3 \exp^{-t/\tau_q} + \Delta_h S_{q_t} \tau_q, \quad (9)$$

504 where the operator  $\Delta_h$  gives the difference of the diabatic flux across the subcloud layer.

505 Table 8 presents the time scales for the SCT cases, based on the average subcloud layer  
506 values from all the LES models. The tendency of the SST was obtained from a linear

507 regression. For all SCT cases the mean value of  $r_{q_t}$  is slightly less than unity, which reflects  
 508 the fact that the subcloud layer is moistening. For the three Composite cases the mean  
 509 depth of the subcloud layer is slightly less than 800 m and the mean horizontal wind speed  
 510 in the subcloud layer  $U_{ml}$  is almost identical. As a result the subcloud layer time scales  $\tau_{CC}$   
 511 and  $\tau_{\theta_v}$  are also very similar.

512 For a sufficiently long simulation time,  $t \gg \tau_{\theta_v}$ , the memory term in the solution for  
 513  $\theta_{v,ml}$ , i.e. the last term in Eq. (8) which includes information about the initial state, vanishes.  
 514 Interestingly, it follows from Eqs. (A-5) and (A-11) that

$$\theta_{v,sfc} - \theta_{v,ml} = (\gamma_{\theta_v} - \Delta_h S_{\theta_v}) \tau_{\theta_v}. \quad (10)$$

515 The constant difference between the subcloud and surface values of  $\theta_v$  has an important  
 516 consequence for the surface buoyancy flux, which according to Eq. (A-3) becomes constant  
 517 in time,

$$\overline{w'\theta'_{v,sfc}} = \frac{(\gamma_{\theta_v} - \Delta_h S_{\theta_v})h}{1 - r_{\theta_v}}. \quad (11)$$

518 The equilibrium surface buoyancy flux value is thus proportional to the depth of the subcloud  
 519 layer and to the horizontal gradient of the sea surface along the path of the air mass. The  
 520 values of the solution for the SCTs are also presented in Table 8. The analytic solutions give  
 521 rather small values for  $\overline{w'\theta'_{v,sfc}}$ , and well explain the behavior of the SHF (see Figs. 3q-t).  
 522 The SHF can be expressed in terms of the surface fluxes of  $\theta_v$  and  $q_t$  as,

$$\text{SHF} \approx \rho c_p (\overline{w'\theta'_{v,sfc}} - \epsilon_I \overline{\theta w'q'_{t,sfc}}), \quad (12)$$

523 with  $\theta$  the potential temperature just above the surface and  $c_p$  the specific heat of dry air.  
 524 For  $\overline{w'\theta'_{v,sfc}} = 0.015 \text{ mKs}^{-1}$  the upper limit of the SHF is about  $16 \text{ Wm}^{-2}$ . The surface  
 525 moisture flux tends to diminish the SHF. For example, if the LHF is  $100 \text{ Wm}^{-2}$  it will lower  
 526 the SHF by about  $7 \text{ Wm}^{-2}$ .

527 With aid of Eqs. (A-3), (A-7), and (A-14) we can express a general solution for the  
 528 surface humidity flux,

$$\overline{w'q'_{t,sfc}} = C_d U_{ml} \left[ \frac{q_{\text{sat},sfc,0} \exp^{t/\tau_{CC}}}{\frac{\tau_{CC}}{\tau_q} + 1} - C_3 \exp^{-t/\tau_q} \right], \quad (13)$$

529 which predicts that  $\overline{w'q'_{t,\text{sfc}}}$  will tend to increase exponentially with time. Substituting the  
 530 mean values from the simulations displayed in Table 8 demonstrates that the analytical  
 531 results for the final hour of the simulations give realistic estimates as compared to the LES  
 532 results. To put the results into perspective, Fig. 13 shows analytical solutions for several  
 533 values of  $r_{q_t}$ . We have neglected the evaporation of drizzle, which for the three Composite  
 534 cases is less than  $1 \text{ Wm}^{-2}$  across the subcloud layer. We used the surface forcing and initial  
 535 conditions from the reference case, in addition to its mean subcloud-layer properties. Because  
 536 the flux ratio  $r_{q_t}$  is a measure of the moisture flux divergence across the subcloud layer, it  
 537 controls the evolution of the moisture in this layer. We notice that its value has a strong  
 538 impact on the evolution of the LHF. Because the LHF is proportional to the difference  
 539 between  $q_{t,\text{ml}}$  and  $q_{\text{sat},\text{sfc},0}$ , a stronger removal of subcloud moisture will trigger a higher  
 540 LHF. Furthermore, we note that  $r_{q_t} = 1$  represents a 'zero-flux divergence' of moisture in  
 541 the subcloud layer, which implies that all the moisture that is evaporated from the surface  
 542 is transported out of the subcloud layer by updrafts. This condition is equivalent to  $q_{t,\text{ml}}$   
 543 being constant in time, which follows directly from  $\tau_q = \infty$  according to Eqs. (A-14) and  
 544 (A-15).

545 In summary, the MLM analysis of the subcloud layer evolution during its Lagrangian  
 546 advection well explains the LES results. For a decoupled boundary layer with a constant  
 547 subcloud layer height, and a fixed value for  $r_{\theta_v}$ , we find that  $\overline{w'\theta'_{v,\text{sfc}}}$  becomes constant in  
 548 time while the surface saturation specific humidity dependency on the SST according to  
 549 Clausius-Clapeyron forces  $\overline{w'q'_{t,\text{sfc}}}$  to grow exponentially in time. An interesting difference  
 550 is found with the first Lagrangian MLM study on stratocumulus by Schubert et al. (1979).  
 551 Their experiment 1 has a similar set up as our subcloud layer MLM analysis, with the SST  
 552 varying linearly in time, and constant values for the wind speed and large-scale divergence.  
 553 For a vertically well-mixed stratocumulus layer they found a gradual increase in the surface  
 554 value of  $\overline{w'\theta'_v}$ .

## 6. Conclusions

Four Lagrangian stratocumulus to shallow cumulus transition experiments were performed with six different LES models. The cases differ predominantly in terms of the amplitude and timescale of the transition. The LES models agree remarkably well in the representation of the evolution of the mean states. For all cases the structure of the boundary layer transforms from a vertically well mixed layer to one in which the subcloud and cloud layers appear as two separated mixed layers, with the stratocumulus layer being warmer and drier as compared to the subcloud layer, a situation which is referred to as decoupling (Nicholls 1984; Bretherton and Wyant 1997). The difference in the thermodynamic state of the subcloud and cloud layers increases for deeper boundary layers, which is found to be in a qualitative agreement with aircraft observations analysed by Wood and Bretherton (2004). The general good agreement between the models in the representation of the boundary-layer evolution can be partly explained by drizzle and solar heating of the cloud layer. Thicker cloud layers such as found for the ASTEX case will produce more precipitation and will absorb more solar radiation during daytime, and vice versa. In this way both processes act to diminish intermodel differences in the LWP. For the Composite cases the earliest timing of the break up of the stratocumulus layer is found for the Fast case, which is predominantly due to a slightly stronger entrainment warming and drying as compared to the Reference and Slow cases.

Superposed to this picture where the boundary layer is deepening due to increasing SSTs, there is a diurnal cycle associated with the absorption of solar radiation within the cloud layer. The models agree well in terms of LWP during the day, but less so in terms of LWP during night. The opposite is true for the cloud cover, which varies considerably among the LES models during daytime. The EULAG model tends to maintain a closed cloud deck which can be attributed to its radiation scheme which gives a somewhat stronger longwave radiative cooling in the cloud layer. The SHF is small and of the order of



581  $10 \text{ Wm}^{-2}$ , whereas the LHF tends to increase with time for all cases.

582 The time evolution of the surface heat fluxes can be well explained by means of a simple  
583 mixed-layer model that is applied to the subcloud layer, and which uses generic bulk features  
584 found from the LES results as boundary conditions. Specifically the model makes use of  
585 the facts that the subcloud layer depth becomes almost constant in time, and that the  
586 buoyancy flux at the top of the subcloud layer tends to approach a fixed negative fraction  
587 of the surface value, similar to what is found for the dry convective boundary layer and  
588 cumulus-topped boundary layers. The critical quantity that controls the magnitude of the  
589 change in the surface evaporation is the moisture flux at the top of the subcloud layer. The  
590 fact that the specific humidity in the subcloud layer increases with time indicates that on  
591 average the surface moisture flux is larger than the value at the top of the subcloud layer.  
592 The LWP budget analysis shows that during periods with stronger turbulence, i.e. during  
593 nighttime, a stronger injection of subcloud layer moisture into the stratocumulus cloud base  
594 is accompanied by a stronger entrainment drying.

595 Fig. 14 presents a schematic of the main findings of the Lagrangian SCTs. The SHF  
596 remains rather small during the equatorwards advection of the air mass, while the LHF  
597 gradually increases. During nighttime the longwave radiative cooling acts to destabilize  
598 the cloud layer, which tends to generate more turbulence and a higher entrainment rate at  
599 the cloud top. Due to stronger turbulence in the cloud layer during the night, subcloud  
600 layer moisture is transported towards the stratocumulus at a rate that exceeds the surface  
601 evaporation during the first night of the three Composite cases, and also during the second  
602 night of the Slow case. This enhanced moisture flux feeds the stratocumulus with liquid  
603 water, thereby competing against the cloud thinning tendency by an increased entrainment  
604 of warm and dry air from just above the inversion. Overall we find that the nocturnal  
605 stratocumulus cloud deck is able to recover from a broken to a closed structure. During  
606 daytime the cloud layer is heated by absorption of solar radiation. This stabilises the cloud  
607 layer with respect to the subcloud layer, which hinders the vertical turbulent transport of

608 subcloud layer moisture to the cloud layer. The warming by the sun, and the reduced  
609 moisture input at the base of the stratocumulus causes it to thin and to break up.

610 The representation of the moisture transport from the top of the subcloud mixed layer  
611 to the stratocumulus layer, and the entrainment of free tropospheric dry air at the top of  
612 the stratocumulus, are essential ingredients to capture the SCT. In fact, in a study on the  
613 representation of the SCT in large-scale models by Neggers (2015) it is found that SCMs  
614 favor a break-up of stratocumulus for inversion conditions that are different to each individual  
615 model. The presence of such modes may be indicative of a local hydrological cycle that is  
616 distinctively different among the models. The finding that the degree of decoupling has  
617 an important consequence for the LWP suggests that the decoupling parameters can be a  
618 helpful quantity in evaluating parameterization schemes for cloud-topped boundary layers  
619 (Dal Gesso et al. 2014). The 3D instantaneous LES (thermo-) dynamic fields may be further  
620 used to evaluate parameterizations used in global models.

621 SCT cases such as discussed here have been simulated to study the effect of changes  
622 in the large-scale forcing conditions in the Hadley cell under climate change conditions  
623 to assess its possible impact of the pace of the transition. For example, Bretherton and  
624 Blossey (2014) investigated and explained the effect of a perturbed radiative forcing, the  
625 overall tropical warming and changes in the inversion stability on the SCT. Likewise, Van  
626 der Dussen et al. (2016) used the LWP budget equation to investigate why a decrease in  
627 the large-scale subsidence extends the lifetime of stratocumulus despite an increase in the  
628 entrainment rate. In addition, both studies investigated the effect of applying a uniform  
629 insolation (constant in time) on the SCTs, which showed that the bulk evolution of the SCT  
630 in terms of boundary-layer deepening is rather similar. Kazil et al. (2015) investigated the  
631 effect of the wind speed on the SCT. They found that a higher wind speed leads to a larger  
632 entrainment rate and a faster growth of the boundary layer, caused by an enhanced buoyant  
633 production of turbulence kinetic energy (TKE) from latent heat release in cloud updrafts.

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## 650 **Appendix A. A mixed-layer model for the subcloud layer**

651 The budget equation for an arbitrary conserved thermodynamic variable  $\psi$  in a horizon-  
652 tally homogeneous atmosphere reads,

$$\frac{d\bar{\psi}}{dt} = -\frac{\partial \overline{w'\psi'}}{\partial z} - \frac{\partial \overline{S_\psi}}{\partial z} \quad (\text{A-1})$$

653 where  $S_\psi$  is a diabatic source term. A vertical integration from the surface to the top of the  
654 subcloud layer  $h$  gives an expression for the vertical mean value  $\psi_{\text{ml}}$ ,

$$\frac{\partial \psi_{\text{ml}}}{\partial t} = \frac{\overline{w'\psi'}_{\text{sfc}} - \overline{w'\psi'}_h}{h} + \frac{\overline{S_{\psi,\text{sfc}}} - \overline{S_{\psi,h}}}{h} \quad (\text{A-2})$$

655 with the subscripts 'sfc' and 'h' denoting the surface and the top of the subcloud layer,  
656 respectively. Because we will apply the budget equation to an air mass that is being advected

657 by the horizontal mean wind, the mean horizontal advection terms can be neglected. The  
 658 vertical advection term due to large-scale subsidence disappears because the assumption of  
 659 well mixedness implies that the vertical gradient of  $\psi$  is zero.

660 In the following we use the notation as presented in Table 7. The surface flux is computed  
 661 from a bulk formula,

$$\overline{w'\psi'}_{\text{sfc}} = C_d U_{\text{ml}} (\psi_{\text{sfc}} - \psi_{\text{ml}}), \quad (\text{A-3})$$

662 with  $C_d = 0.0012$  a bulk drag coefficient and  $U_{\text{ml}}$  the absolute value of the mean horizontal  
 663 wind speed in the subcloud layer. To obtain analytical solutions we will assume that the sea  
 664 surface temperature increases linearly with time,

$$\text{SST}(t) = \text{SST}_0 + \gamma_T t. \quad (\text{A-4})$$

665 Likewise we can express the surface virtual potential temperature as

$$\theta_{\text{v,sfc}}(t) = \theta_{\text{v,sfc},0} + \gamma_{\theta_v} t. \quad (\text{A-5})$$

666 Since the change in  $\theta_{\text{v,sfc}}$  is dominated by changes in the SST we will approximate  $\gamma_{\theta_v} \approx$   
 667  $\gamma_T(1 + \epsilon_I q_{\text{sat,sfc},0})/\Pi$ , with  $\Pi$  the Exner function, and  $\epsilon_I \approx 0.608$ .

668 To compute the temporal variation of the surface moisture flux we will use an approxi-  
 669 mated form of the Clausius-Clapeyron equation (Stevens 2006),

$$q_{\text{sat,sfc}}(\text{SST}) = q_{\text{sat,sfc},0} \exp \left[ \frac{L_v}{R_v \text{SST}_0^2} (\text{SST} - \text{SST}_0) \right]. \quad (\text{A-6})$$

670 For a linear increase of the temperature with time according to Eq. (A-4),  $q_{\text{sat,sfc}}$  will increase  
 671 exponentially with time,

$$q_{\text{sat,sfc}}(t) = q_{\text{sat,sfc},0} e^{t/\tau_{\text{CC}}}. \quad (\text{A-7})$$

672 Given this framework, the tendency for the virtual potential temperature in the subcloud  
 673 layer  $\theta_{\text{v,ml}}$  is governed by the turbulent flux divergence which can be expressed in terms of  
 674 the flux ratio  $r_{\theta_v}$ ,

$$\frac{\partial \theta_{\text{v,ml}}}{\partial t} = (1 - r_{\theta_v}) \frac{\overline{w'\theta'_{\text{v,sfc}}}}{h} + \Delta_h S_{\theta_v}, \quad (\text{A-8})$$

675 with the source term representing the divergence of the net radiative flux. For the Composite  
676 SCTs the net longwave radiative flux varies between 1 and 2  $\text{Wm}^{-2}$  during nighttime and  
677 daytime, respectively, across a vertical layer of 100 m below the clouds. The maximum solar  
678 radiative flux divergence is about 3  $\text{Wm}^{-2}$  per 100 m, which leaves a negligibly small diurnal  
679 mean radiative forcing of the subcloud layer.

680 On the basis of the results presented in Fig. 3, we will ignore variations of  $h$  in time.  
681 In addition, we take  $\Delta_h S_{\theta_v}$  constant with time. Using Eqs. (A-3) and (A-5), this allows to  
682 express Eq. (A-8) as,

$$\frac{\partial \theta_{v,\text{ml}}}{\partial t} = \frac{\theta_{v,\text{sfc},0} + \gamma_{\theta_v} t - \theta_{v,\text{ml}}}{\tau_{\theta_v}} + \Delta_h S_{\theta_v}. \quad (\text{A-9})$$

683 The solution of Eq. (A-9) is given by

$$\theta_{v,\text{ml}}(t) = \gamma_{\theta_v} t + C_1 + C_2 \exp^{-t/\tau_{\theta_v}}, \quad (\text{A-10})$$

684 with

$$C_1 = \theta_{v,\text{sfc},0} - \gamma_{\theta_v} \tau_{\theta_v} + \Delta_h S_{\theta_v} \tau_{\theta_v}. \quad (\text{A-11})$$

685 The constant  $C_2$  follows from the initial condition,

$$C_2 = \theta_{v,\text{ml},0} - C_1 = \theta_{v,\text{ml},0} - \theta_{v,\text{sfc},0} + \gamma_{\theta_v} \tau_{\theta_v} - \Delta_h S_{\theta_v} \tau_{\theta_v}. \quad (\text{A-12})$$

686 The budget equation for  $q_{t,\text{ml}}$  can be written as

$$\frac{\partial q_{t,\text{ml}}}{\partial t} = -\frac{q_{t,\text{ml}}}{\tau_q} + \frac{q_{\text{sat},\text{sfc},0}}{\tau_q} e^{t/\tau_{\text{CC}}} + \Delta_h S_{q_t}. \quad (\text{A-13})$$

687 The term  $\Delta_h S_{q_t}$  represents the amount of rain water that evaporates in the subcloud  
688 layer which we take constant in time. To allow for an analytical solution we will neglect  
689 diurnal variations in  $r_{q_t}$ , which gives a solution of the following form,

$$q_{t,\text{ml}}(t) = \frac{q_{\text{sat},\text{sfc},0}}{1 + \frac{\tau_q}{\tau_{\text{CC}}}} \exp^{t/\tau_{\text{CC}}} + C_3 \exp^{-t/\tau_q} + \Delta_h S_{q_t} \tau_q, \quad (\text{A-14})$$

690 with

$$C_3 = q_{t,\text{ml},0} - \frac{q_{\text{sat},\text{sfc},0}}{1 + \frac{\tau_q}{\tau_{\text{CC}}}} - \Delta_h S_{q_t} \tau_q. \quad (\text{A-15})$$

691

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865		except for the surface fluxes which represent the analytical results at the end	
866		of the simulations.	44

		ASTEX	Fast	Reference	Slow
$p_{\text{sfc}}$	hPa	1029.0	1015.9	1016.8	1017.6
latitude	$^{\circ}\text{N}$	34	25	25	25
longitude	$^{\circ}\text{W}$	25	125	125	125
date		13 June	15 July	15 July	15 July
Div	$10^{-6} \text{ s}^{-1}$	-	1.9	1.86	1.84
$z_{\text{Div}}$	km	1.6	2	2	2

TABLE 1. Details of the simulations. Div represents the large-scale divergence of the horizontal mean wind velocities, which is constant in time and constant up to a height of  $z_{\text{Div}}$ , except for the ASTEX case in which the divergence varies with time.

LES model	Institute	References	Participants
DALES	TU Delft, The Netherlands	Heus et al. (2010)	Van der Dussen
MPI/UCLA	MPI-Hamburg, Germany	Stevens et al. (2005b)	Sandu
SAM	U. Washington, USA	Khairoutdinov and Randall (2005)	Blossey
MOLEM	UKMO, UK	Lock (2009)	Lock
DHARMA	NASA GISS, USA	Vogelmann et al. (2015)	Ackerman
EULAG	U. Warsaw, Poland	Prusa et al. (2008)	Jarecka

TABLE 2. Participating models and contributors.

	D1	N1	D2	N2	D3	N3
ASTEX		0–6	7–21	22–30	31–40	
Composite Cases	0–9	10–19	20–33	34–43	44–57	58–67

TABLE 3. Summary of periods of daytime (D1, D2 and D3) and nighttime (N1, N2 and N3), respectively, and the corresponding begin and end times in hours from the start of the simulations.



		Time	ASTEX	Fast	Reference	Slow
$z_{\text{cu,base}}$	m	D1		$560 \pm 30$	$530 \pm 30$	$580 \pm 20$
		N1	$260 \pm 20$	$710 \pm 40$	$720 \pm 20$	$670 \pm 20$
		D2	$360 \pm 10$	$770 \pm 30$	$790 \pm 10$	$740 \pm 20$
		N2	$500 \pm 20$	$750 \pm 40$	$820 \pm 30$	$880 \pm 30$
		D3	$526 \pm 9$	$810 \pm 50$	$870 \pm 20$	$890 \pm 30$
		N3		$800 \pm 50$	$840 \pm 30$	$890 \pm 30$
$z_i$	m	D1		$1038 \pm 6$	$968 \pm 7$	$902 \pm 6$
		N1	$770 \pm 20$	$1260 \pm 20$	$1120 \pm 10$	$1010 \pm 20$
		D2	$1060 \pm 50$	$1520 \pm 50$	$1310 \pm 30$	$1170 \pm 30$
		N2	$1480 \pm 50$	$1700 \pm 100$	$1480 \pm 60$	$1320 \pm 40$
		D3	$1770 \pm 60$	$1900 \pm 100$	$1650 \pm 80$	$1470 \pm 50$
		N3		$2100 \pm 200$	$1800 \pm 100$	$1600 \pm 80$
LWP	$\text{gm}^{-2}$	D1		$30 \pm 10$	$30 \pm 10$	$51 \pm 7$
		N1	$210 \pm 20$	$80 \pm 30$	$90 \pm 20$	$80 \pm 10$
		D2	$130 \pm 20$	$30 \pm 20$	$40 \pm 10$	$50 \pm 10$
		N2	$80 \pm 20$	$50 \pm 30$	$50 \pm 20$	$90 \pm 30$
		D3	$30 \pm 10$	$30 \pm 20$	$30 \pm 10$	$40 \pm 20$
		N3		$30 \pm 20$	$40 \pm 30$	$40 \pm 20$
cc	0-1	D1		$0.98 \pm 0.03$	$0.982 \pm 0.009$	$0.9989 \pm 0.0005$
		N1	$1.0 \pm 0.0$	$0.99 \pm 0.02$	$0.998 \pm 0.001$	$0.9991 \pm 0.0004$
		D2	$0.9994 \pm 0.0004$	$0.9 \pm 0.2$	$0.95 \pm 0.03$	$0.98 \pm 0.01$
		N2	$0.996 \pm 0.005$	$0.9 \pm 0.1$	$0.98 \pm 0.02$	$0.997 \pm 0.002$
		D3	$0.8 \pm 0.2$	$0.7 \pm 0.2$	$0.89 \pm 0.08$	$0.96 \pm 0.02$
		N3		$0.8 \pm 0.2$	$0.93 \pm 0.06$	$0.97 \pm 0.02$
$w_e$	$\text{cm s}^{-1}$	D1		$0.65 \pm 0.06$	$0.49 \pm 0.04$	$0.39 \pm 0.04$
		N1	$1.1 \pm 0.2$	$1.07 \pm 0.08$	$0.80 \pm 0.05$	$0.62 \pm 0.05$
		D2	$1.21 \pm 0.08$	$0.62 \pm 0.09$	$0.52 \pm 0.05$	$0.49 \pm 0.04$
		N2	$1.49 \pm 0.04$	$1.0 \pm 0.1$	$0.83 \pm 0.05$	$0.74 \pm 0.02$
		D3	$0.71 \pm 0.08$	$0.6 \pm 0.1$	$0.54 \pm 0.07$	$0.47 \pm 0.05$
		N3		$0.9 \pm 0.1$	$0.9 \pm 0.1$	$0.79 \pm 0.08$
SHF	$\text{Wm}^{-2}$	D1		$11 \pm 1$	$7.1 \pm 0.9$	$11 \pm 1$
		N1	$7 \pm 1$	$9 \pm 1$	$6.2 \pm 0.7$	$10 \pm 1$
		D2	$14 \pm 1$	$7 \pm 1$	$6.6 \pm 0.7$	$9 \pm 1$
		N2	$5.6 \pm 0.8$	$9 \pm 2$	$7 \pm 1$	$7 \pm 1$
		D3	$2.1 \pm 0.2$	$8 \pm 2$	$8 \pm 1$	$6.6 \pm 0.9$
		N3		$8 \pm 2$	$9 \pm 2$	$8 \pm 1$
LHF	$\text{Wm}^{-2}$	D1		$104 \pm 7$	$80 \pm 3$	$90 \pm 4$
		N1	$60 \pm 10$	$126 \pm 7$	$105 \pm 4$	$103 \pm 5$
		D2	$100 \pm 10$	$138 \pm 7$	$119 \pm 6$	$110 \pm 6$
		N2	$94 \pm 7$	$151 \pm 7$	$130 \pm 6$	$121 \pm 6$
		D3	$56 \pm 6$	$167 \pm 9$	$153 \pm 8$	$133 \pm 8$
		N3		$169 \pm 8$	$159 \pm 8$	$150 \pm 10$
$\text{SW}_{\text{net,sfc}}$	$\text{Wm}^{-2}$	D1		$-410 \pm 30$	$-420 \pm 40$	$-350 \pm 30$
		N1	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$
		D2	$-230 \pm 30$	$-440 \pm 60$	$-410 \pm 40$	$-380 \pm 40$
		N2	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$
		D3	$-510 \pm 80$	$-470 \pm 80$	$-450 \pm 60$	$-420 \pm 50$
		N3	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$	$0 \pm 0$
$\text{LW}_{\text{net,sfc}}$	$\text{Wm}^{-2}$	D1		$30 \pm 6$	$28 \pm 3$	$23 \pm 2$
		N1	$11.0 \pm 0.7$	$27 \pm 5$	$22.5 \pm 0.9$	$22 \pm 2$
		D2	$19 \pm 2$	$40 \pm 10$	$36 \pm 4$	$31 \pm 2$
		N2	$24 \pm 2$	$40 \pm 10$	$33 \pm 4$	$27.0 \pm 0.9$
		D3	$40 \pm 10$	$50 \pm 10$	$46 \pm 7$	$39 \pm 4$
		N3		$50 \pm 10$	$41 \pm 7$	$37 \pm 4$

TABLE 4. Mean values and their standard deviations during the daytime and nighttime periods according to Table 3. The standard deviation is rounded to one significant digit. However, for compact notation we express, for example,  $(10 \pm 2) \cdot 10^1$  as  $100 \pm 20$ .

	Time	ASTEX	Fast	Reference	Slow
$r_{\theta_v}$	D1		$-0.19 \pm 0.08$	$-0.2 \pm 0.1$	$0.1 \pm 0.2$
	N1	$-0.2 \pm 0.4$	$-0.3 \pm 0.1$	$0.0 \pm 0.4$	$0.4 \pm 0.3$
	D2	$-0.20 \pm 0.07$	$-0.17 \pm 0.05$	$-0.20 \pm 0.07$	$-0.2 \pm 0.1$
	N2	$-0.20 \pm 0.09$	$-0.17 \pm 0.07$	$-0.21 \pm 0.05$	$-0.1 \pm 0.2$
	D3	$-0.09 \pm 0.09$	$-0.14 \pm 0.07$	$-0.16 \pm 0.04$	$-0.19 \pm 0.06$
	N3		$-0.16 \pm 0.08$	$-0.17 \pm 0.07$	$-0.21 \pm 0.06$
	D1		$0.87 \pm 0.07$	$1.1 \pm 0.1$	$0.93 \pm 0.05$
$r_{q_t}$	N1	$1.09 \pm 0.09$	$1.04 \pm 0.07$	$1.26 \pm 0.08$	$1.09 \pm 0.05$
	D2	$0.83 \pm 0.07$	$0.67 \pm 0.07$	$0.72 \pm 0.05$	$0.84 \pm 0.08$
	N2	$0.92 \pm 0.05$	$0.95 \pm 0.03$	$0.97 \pm 0.05$	$1.1 \pm 0.1$
	D3	$0.7 \pm 0.2$	$0.79 \pm 0.05$	$0.70 \pm 0.07$	$0.67 \pm 0.05$
	N3		$0.83 \pm 0.07$	$0.91 \pm 0.02$	$0.92 \pm 0.07$

TABLE 5. Mean values of the flux ratios  $r_{\theta_v}$  and  $r_{q_t}$  and their standard deviations during the daytime and nighttime periods according to Table 3.

$\partial\text{LWP}/\partial t$ ( $\text{gm}^{-2}\text{s}^{-1}$ )	Time	ASTEX	Fast	Reference	Slow	
Rad <sub>LW</sub>	$\frac{\eta\gamma}{c_p} [\text{LW}_{\text{net}}(z_t) - \text{LW}_{\text{net}}(z_b)]$	D1		$54 \pm 9$	$54 \pm 9$	$57 \pm 6$
		N1	$62 \pm 1$	$60 \pm 8$	$61 \pm 5$	$59 \pm 6$
		D2	$65 \pm 1$	$50 \pm 10$	$50 \pm 10$	$53 \pm 9$
		N2	$67 \pm 4$	$59 \pm 7$	$56 \pm 8$	$60 \pm 6$
		D3	$57 \pm 7$	$60 \pm 10$	$50 \pm 10$	$50 \pm 10$
		N3		$50 \pm 10$	$50 \pm 10$	$56 \pm 9$
Rad <sub>SW</sub>	$\frac{\eta\gamma}{c_p} [\text{SW}_{\text{net}}(z_t) - \text{SW}_{\text{net}}(z_b)]$	D1		$-20 \pm 4$	$-19 \pm 3$	$-25 \pm 3$
		N1	$0.0 \pm 0.0$	$-0.0 \pm 0.0$	$-0.0 \pm 0.0$	$-0.0 \pm 0.0$
		D2	$-34 \pm 2$	$-16 \pm 4$	$-17 \pm 4$	$-20 \pm 4$
		N2	$0.0 \pm 0.0$	$-0.0 \pm 0.0$	$-0.0 \pm 0.0$	$-0.0 \pm 0.0$
		D3	$-19 \pm 1$	$-19 \pm 6$	$-14 \pm 4$	$-16 \pm 4$
		N3		$-0.0 \pm 0.0$	$-0.0 \pm 0.0$	$-0.0 \pm 0.0$
Ent <sub>heat</sub>	$-\rho w_e \Pi \gamma \eta \Delta \theta_1$	D1		$-51 \pm 6$	$-45 \pm 5$	$-41 \pm 5$
		N1	$-70 \pm 10$	$-87 \pm 8$	$-74 \pm 5$	$-63 \pm 5$
		D2	$-70 \pm 6$	$-52 \pm 6$	$-46 \pm 5$	$-48 \pm 4$
		N2	$-93 \pm 5$	$-89 \pm 7$	$-74 \pm 7$	$-72 \pm 3$
		D3	$-43 \pm 5$	$-58 \pm 8$	$-48 \pm 5$	$-44 \pm 5$
		N3		$-80 \pm 10$	$-80 \pm 10$	$-80 \pm 10$
Ent <sub>dry</sub>	$\rho w_e \eta \Delta q_t$	D1		$-44 \pm 5$	$-44 \pm 5$	$-38 \pm 4$
		N1	$-28 \pm 6$	$-76 \pm 6$	$-72 \pm 4$	$-60 \pm 4$
		D2	$-48 \pm 5$	$-45 \pm 5$	$-45 \pm 5$	$-47 \pm 4$
		N2	$-83 \pm 4$	$-80 \pm 10$	$-70 \pm 5$	$-71 \pm 2$
		D3	$-44 \pm 6$	$-53 \pm 6$	$-45 \pm 5$	$-44 \pm 4$
		N3		$-77 \pm 7$	$-70 \pm 10$	$-74 \pm 8$
Base <sub>heat</sub>	$-\rho \eta \Pi \gamma \overline{w' \theta'_1}(z_b)$	D1		$11 \pm 2$	$7 \pm 2$	$4 \pm 3$
		N1	$10 \pm 7$	$17 \pm 4$	$8 \pm 5$	$0 \pm 4$
		D2	$21 \pm 4$	$10 \pm 2$	$8 \pm 2$	$8 \pm 2$
		N2	$16 \pm 2$	$22 \pm 2$	$13 \pm 2$	$9 \pm 3$
		D3	$0 \pm 1$	$15.0 \pm 0.6$	$12 \pm 3$	$7 \pm 2$
		N3		$20 \pm 2$	$20.2 \pm 0.8$	$14 \pm 1$
Base <sub>moist</sub>	$\rho \eta \overline{w' q'_t}(z_b)$	D1		$45 \pm 4$	$47 \pm 6$	$43 \pm 2$
		N1	$37 \pm 4$	$71 \pm 7$	$70 \pm 4$	$61 \pm 4$
		D2	$44 \pm 3$	$43 \pm 4$	$44 \pm 6$	$48 \pm 4$
		N2	$62 \pm 3$	$70 \pm 10$	$69 \pm 5$	$73 \pm 6$
		D3	$34 \pm 2$	$53 \pm 6$	$46 \pm 4$	$44 \pm 5$
		N3		$73 \pm 4$	$70 \pm 10$	$72 \pm 8$
Prec	$-\rho [P(z_t) - P(z_b)]$	D1		$-1 \pm 1$	$-1 \pm 1$	$-1 \pm 1$
		N1	$-40 \pm 30$	$-3 \pm 2$	$-3 \pm 2$	$-2 \pm 1$
		D2	$-18 \pm 6$	$-2 \pm 2$	$-2 \pm 1$	$-2 \pm 1$
		N2	$-4 \pm 2$	$-3 \pm 3$	$-2 \pm 2$	$-3 \pm 2$
		D3	$-1.3 \pm 0.4$	$-2 \pm 1$	$-2 \pm 2$	$-2 \pm 2$
		N3		$-4 \pm 4$	$-3 \pm 2$	$-2 \pm 2$
Ent <sub>z<sub>i</sub></sub>	$-\rho w_e h_{\text{cld}} \Gamma_{q_1}$	D1		$9 \pm 2$	$7 \pm 2$	$7.1 \pm 0.6$
		N1	$41 \pm 7$	$23 \pm 5$	$18 \pm 2$	$15 \pm 1$
		D2	$33 \pm 4$	$10 \pm 2$	$9 \pm 2$	$9 \pm 1$
		N2	$33 \pm 4$	$18 \pm 6$	$14 \pm 3$	$17 \pm 3$
		D3	$14 \pm 2$	$11 \pm 4$	$8 \pm 2$	$8 \pm 2$
		N3		$15 \pm 5$	$14 \pm 6$	$13 \pm 4$

TABLE 6. Mean values and their standard deviations for some key LWP budget terms according to Eq. (4) during the daytime and nighttime periods according to Table 3.

	$SST_0$
initial surface conditions	$\theta_{v,\text{sfc},0}$
	$q_{\text{sat},\text{sfc},0}$
initial mixed layer conditions	$\theta_{v,\text{ml},0}$
	$q_{t,\text{ml},0}$
surface boundary conditions	$\gamma_T \equiv \partial SST / \partial t$
	$\gamma_{\theta_v} \equiv \partial \theta_{v,\text{sfc}} / \partial t$
	$\tau_{\theta_v} \equiv \frac{h}{(1 - r_{\theta_v}) C_d U_{\text{ml}}}$
time scales	$\tau_q \equiv \frac{h}{(1 - r_{q_t}) C_d U_{\text{ml}}}$
	$\tau_{\text{CC}} = \frac{R_v S_{\text{T}}^2}{L_v \gamma_T}$
	$C_1 = \theta_{v,\text{sfc},0} - \tau_{\theta_v} (\gamma_{\theta_v} - \Delta_h S_{\theta_v})$
constants	$C_2 = \theta_{v,\text{ml},0} - \theta_{v,\text{sfc},0} + \tau_{\theta_v} (\gamma_{\theta_v} - \Delta_h S_{\theta_v})$
	$C_3 = q_{t,\text{ml},0} - \frac{q_{\text{sat},\text{sfc},0}}{1 + \frac{\tau_q}{\tau_{\text{CC}}}} - \Delta_h S_{q_t} \tau_q$

TABLE 7. Summary of the boundary conditions used for the subcloud mixed layer model, its time scales, and the definitions of the constants  $C_1$ ,  $C_2$  and  $C_3$ .

		Fast	Reference	Slow
$\gamma_T$	$\text{K day}^{-1}$	1.7	1.9	1.7
$r_{q_t}$		0.8	0.9	0.9
$U_{\text{ml}}$	$\text{ms}^{-1}$	5.7	5.3	5.1
$h$	m	756	781	789
$\tau_{\theta_v}$	h	25.6	28.7	30.0
$\tau_q$	h	152	296	318
$\tau_{\text{CC}}$	h	227	202	218
$\overline{w'\theta'_v}$ , Eq. (11)	$\text{mKs}^{-1}$	0.012	0.014	0.013
LHF, Eq. (13)	$\text{Wm}^{-2}$	165	170	162

TABLE 8. Average values as obtained during the entire run and from all the LES models, except for the surface fluxes which represent the analytical results at the end of the simulations.

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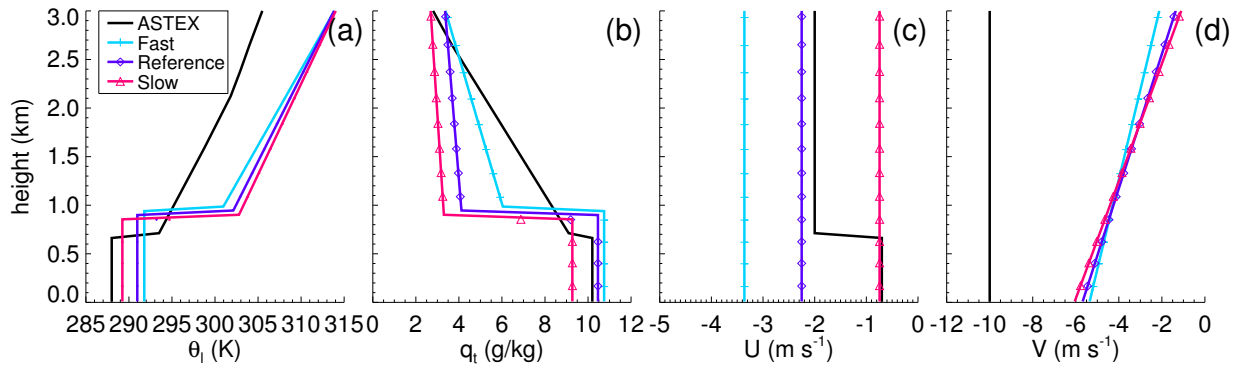


FIG. 1. Initial vertical profiles of (a) the liquid water potential temperature  $\theta_l$ , (b) the total water specific humidity  $q_t$ , and the horizontal wind velocity components (c)  $U$  and (d)  $V$  for the ASTEX, Fast, Reference and Slow cases. The line styles are according to the legend.

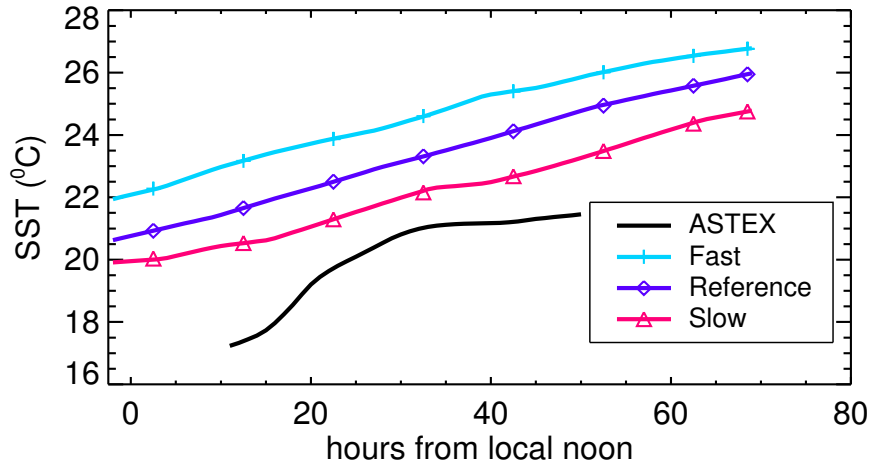


FIG. 2. Prescribed sea surface temperature for the ASTEX, Fast, Reference and Slow cases. The line styles are according to the legend.

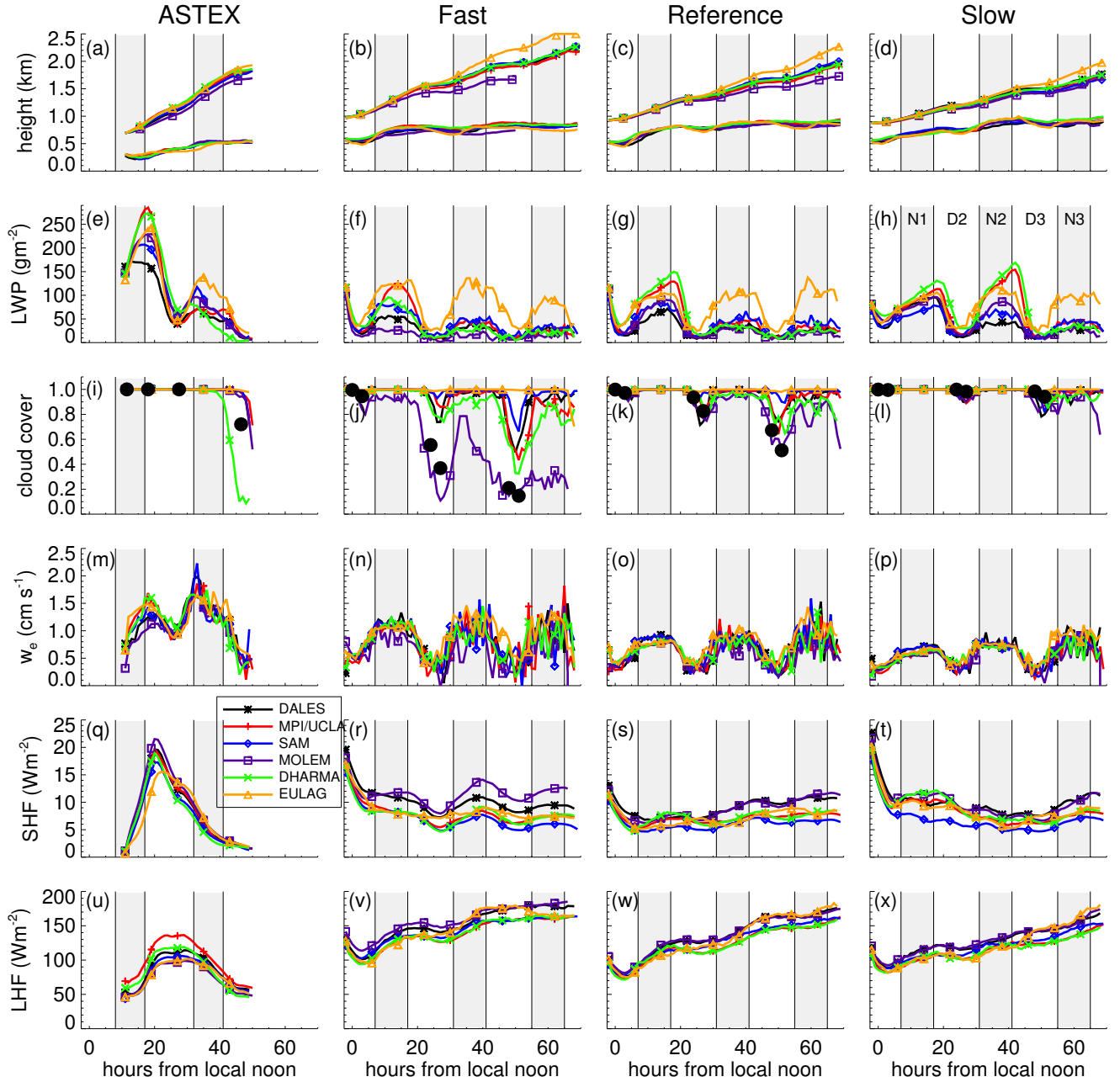


FIG. 3. Time series of the (a-d) lowest cumulus cloud base height (lower solid lines without symbols) and the mean inversion height (upper solid lines with symbols), (e-h) the domain-averaged liquid water path LWP, (i-l) the cloud cover, (m-p) the entrainment velocity  $w_e$ , (q-t) the sensible heat flux SHF and (u-x) the latent heat flux LHF. From left to right the columns present results of the ASTEX, Fast, Reference and Slow cases, respectively. The line styles are according to the legend displayed in figure q. The filled black big circles in i indicate the cloud cover as derived from aircraft observations, and in j,k,l they represent retrievals from the MODIS satellite along the trajectories of the Composite cases and can be considered as an upper bound of the real cloud fraction [see Appendix A of Sandu et al. (2010)]. The grey shaded bands indicate periods of nighttime (denoted at the top of Fig. 3h) according to Table 3.

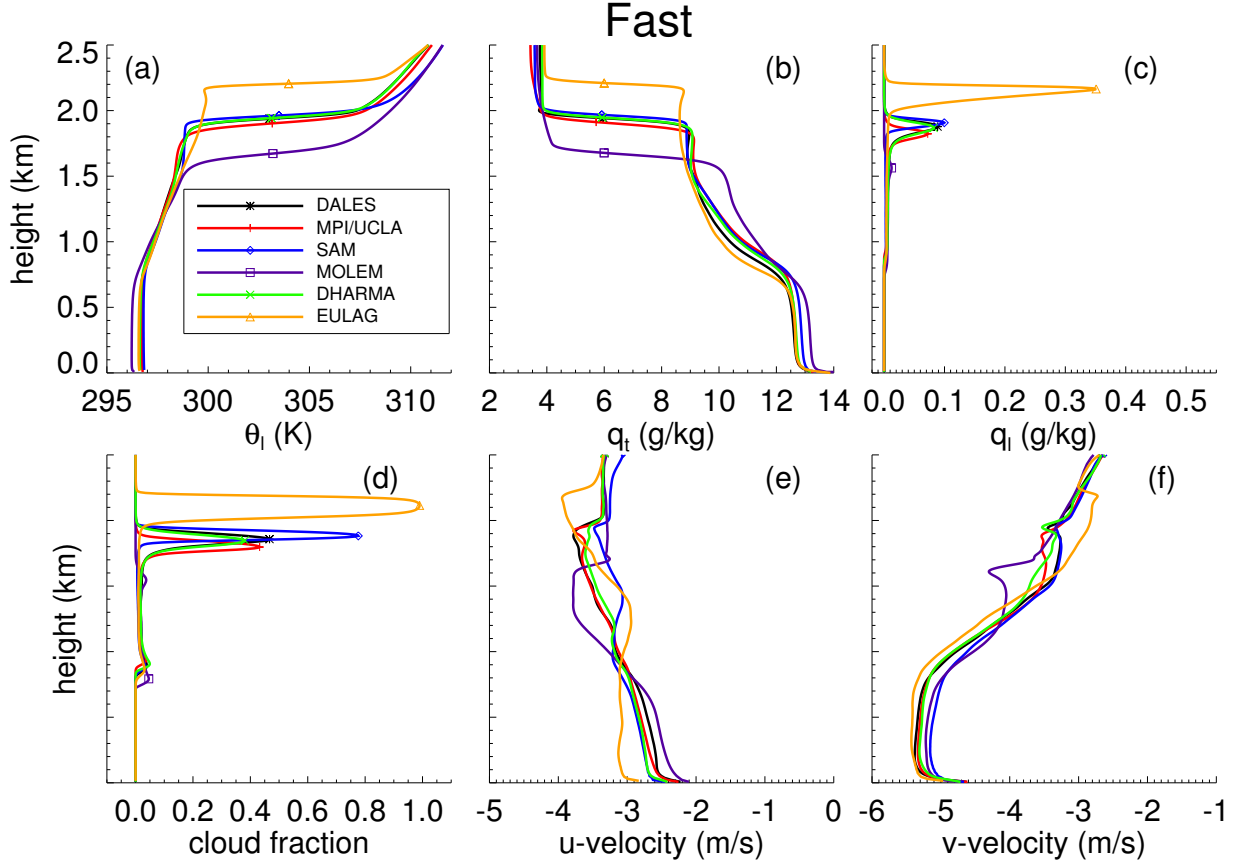


FIG. 4. Vertical profiles of (a) the liquid water potential temperature  $\theta_l$ , (b) the total specific humidity  $q_t$ , (c) the liquid water specific humidity  $q_l$ , (d) the cloud fraction, (e) the east-west velocity  $U$ , and (f) the north-south velocity  $V$  for the Fast case. The lines represent hourly-mean horizontally slab averaged values obtained during the 48th hour from local noon. The line styles are according to the legend.

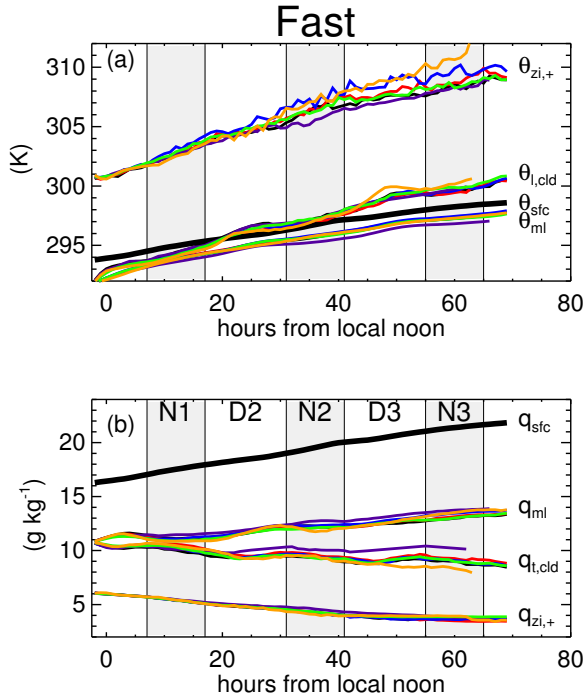


FIG. 5. Evolution of (a) the (liquid water) potential temperature and (b) the (total) specific humidity just above the inversion ( $z_i^+$ ), and their vertical mean values in the subcloud (sub) and the stratocumulus cloud layer (cld) for the Fast case. For easy reference the prescribed values at the surface (sfc) are also plotted. The line styles are according to the legend as in Fig. 4a. The grey shaded bands indicate nighttime periods according to Table 3.

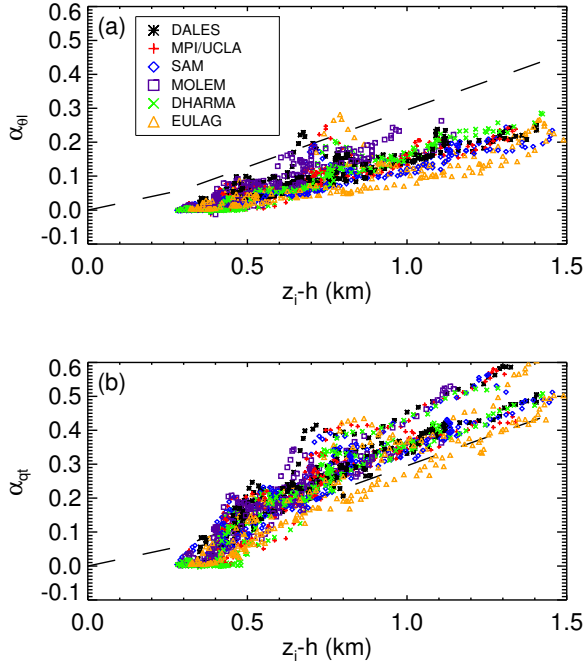


FIG. 6. The decoupling parameters (a)  $\alpha_{\theta_1}$  and (b)  $\alpha_{q_t}$  as a function of the cloud layer depth. The dashed lines indicate a fit using the aircraft observations of  $\alpha_{q_t}$  presented by Wood and Bretherton (2004), their Fig. 5. The symbols are according to the legend.

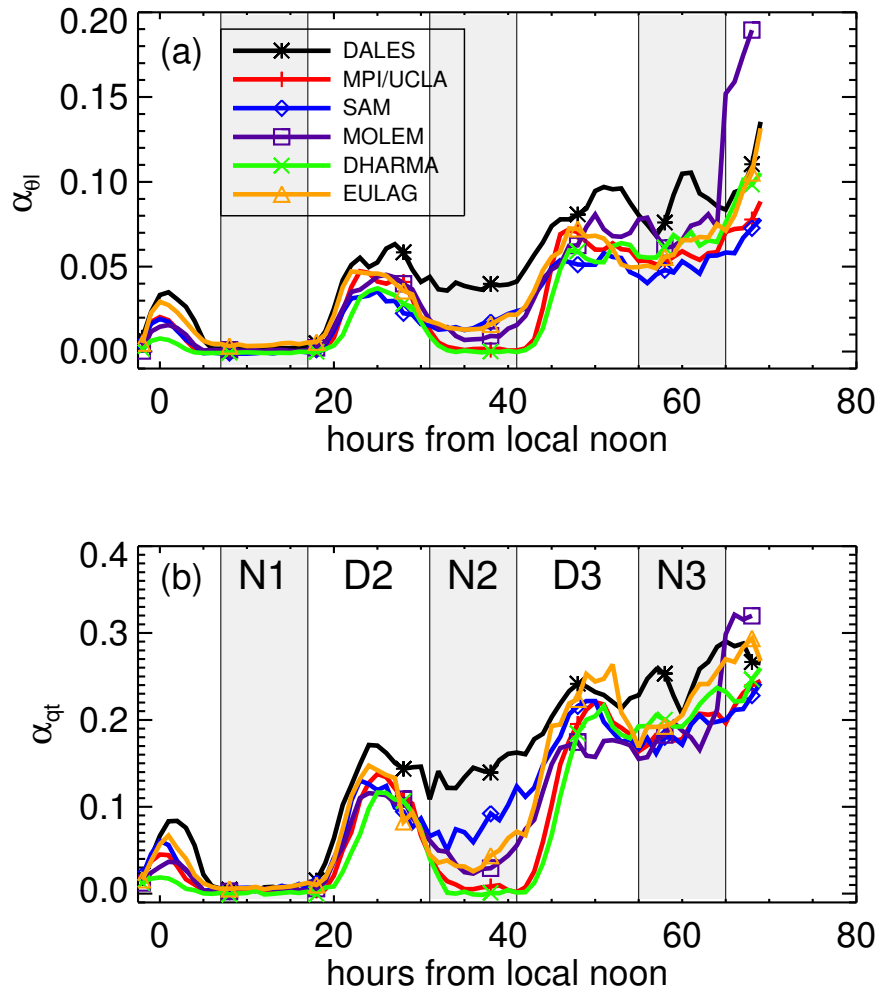


FIG. 7. The time evolution of the decoupling parameters of (a-d)  $\alpha_{\theta_1}$  and (e-h)  $\alpha_{q_t}$  for the Slow case. The line colors and symbols are according to the legend. The grey shaded bands indicate nighttime periods according to Table 3.

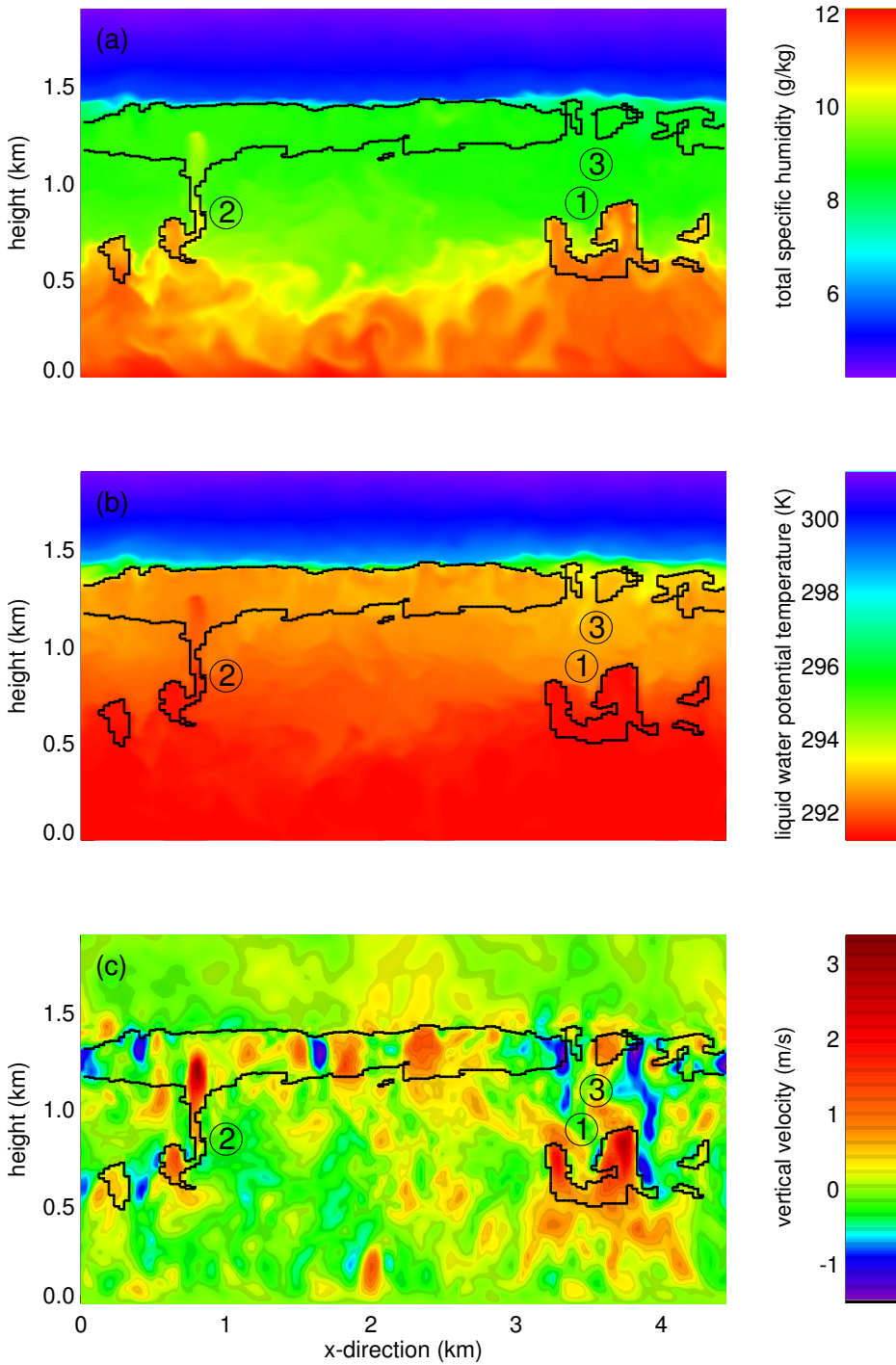


FIG. 8. Instantaneous fields in the vertical plane for (a) the total water specific humidity, (b) the liquid water potential temperature, and (c) the vertical velocity as obtained 36 hours from local noon from the DALES ASTEX run. The thick solid black lines indicate the contours of the cloud edges. See text for an explanation of the areas that are indicated by the encircled numbers.



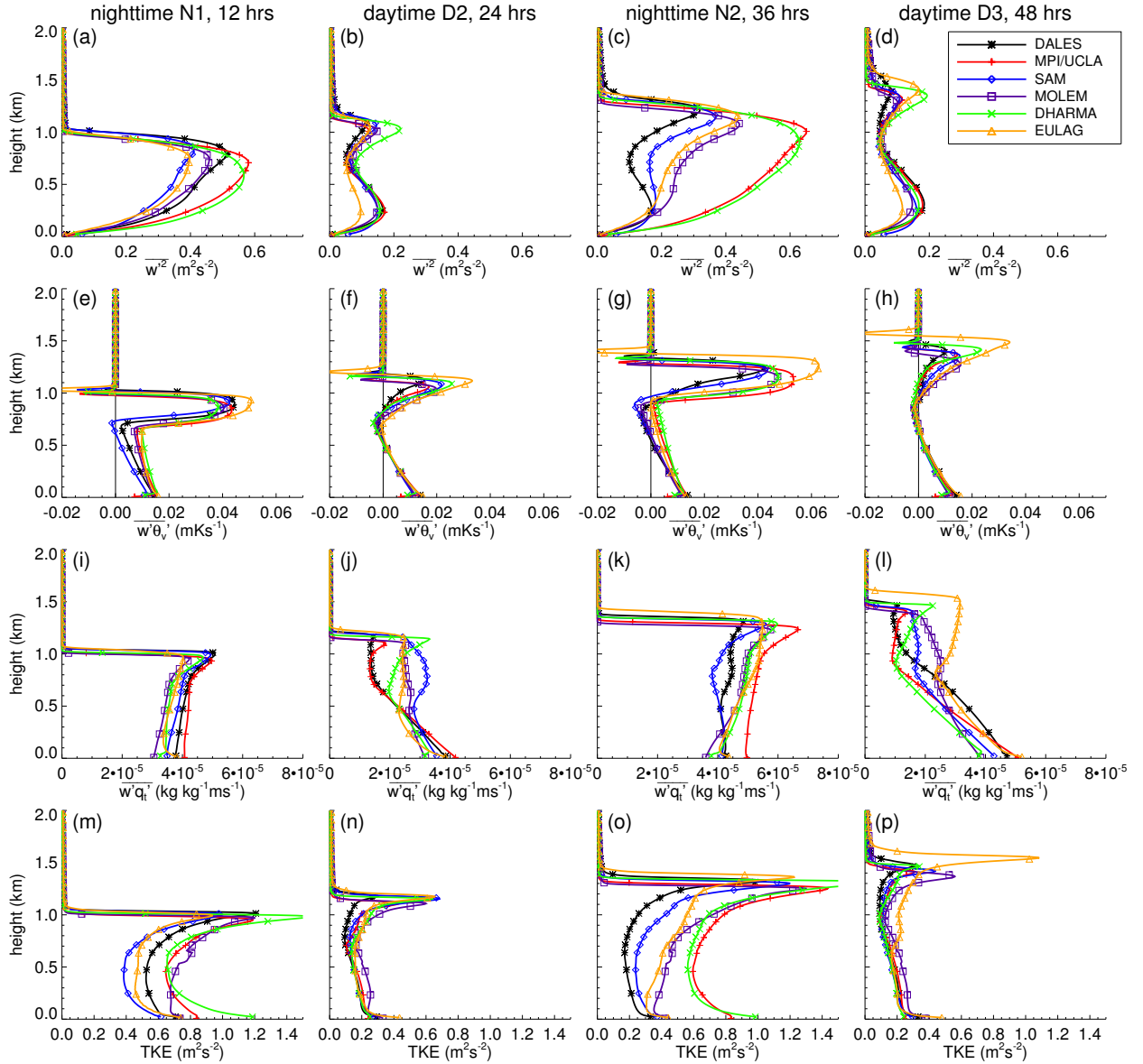


FIG. 9. Hourly-mean turbulence statistics for the 'Slow' case at four selected times. The profiles at 12 and 36 hrs from local noon are at midnight, and 24 and 48 hrs represent conditions during local noon. The first row (a-d) shows the vertical velocity variance, the second row (e-h) the virtual potential temperature flux, the third row (i-l) the total water specific humidity flux and the bottom row (m-p) the turbulent kinetic energy. The line colors and symbols are as in the legend shown in figure d. The thin black vertical line in the plots showing the virtual potential temperature flux indicates a zero value for easy reference.

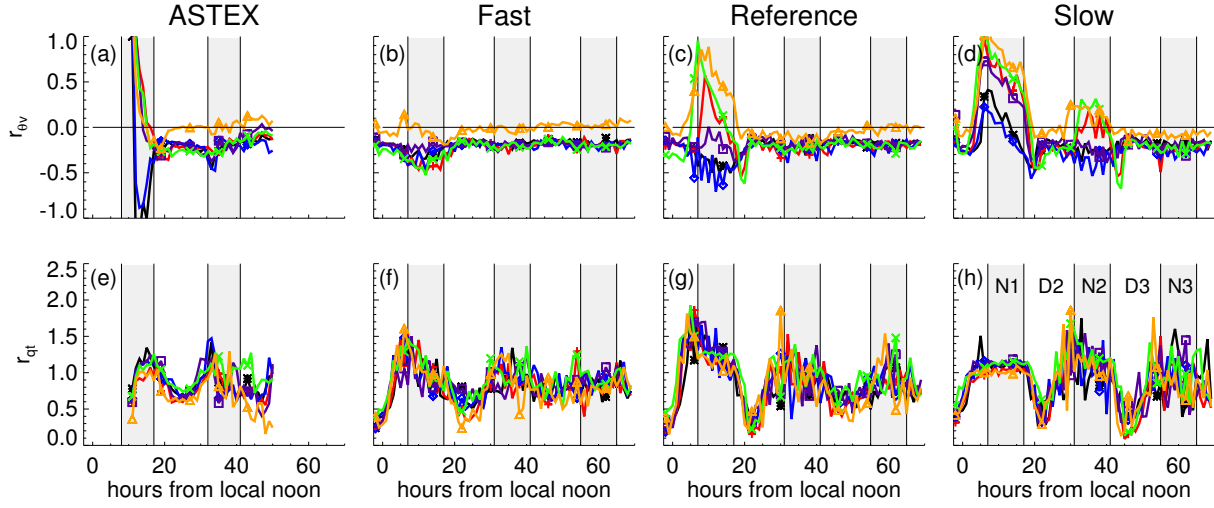


FIG. 10. The time evolution of the flux ratios for the (a-d) buoyancy ( $r_{\theta_v}$ ) and (e-h) the total water specific humidity ( $r_{q_t}$ ) computed according to Eq. (2). The line styles are according to the legend. The grey shaded bands indicate nighttime periods according to Table 3. The line colors and symbols are as in the legend shown in Fig. 4a. The thin solid black line in a-d represents the zero line.

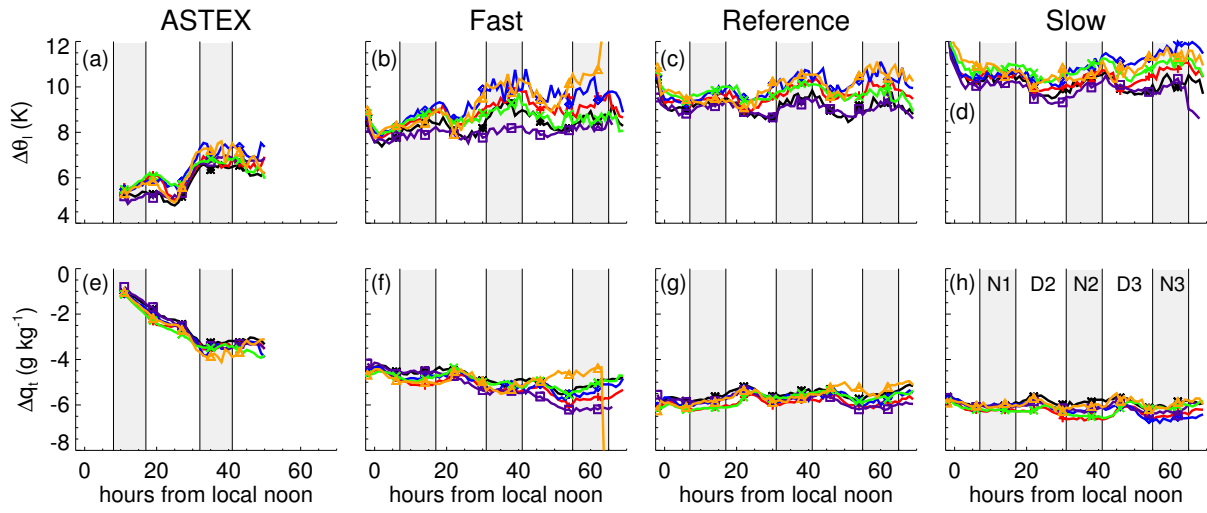


FIG. 11. Time evolution of the inversion jumps of (a-d)  $\Delta\theta_1$  and (e-h)  $\Delta q_t$ . The line colors and symbols are as in the legend shown in Fig. 4a. The grey shaded bands indicate nighttime periods according to Table 3. The line colors and symbols are as in the legend shown in Fig. 4a.

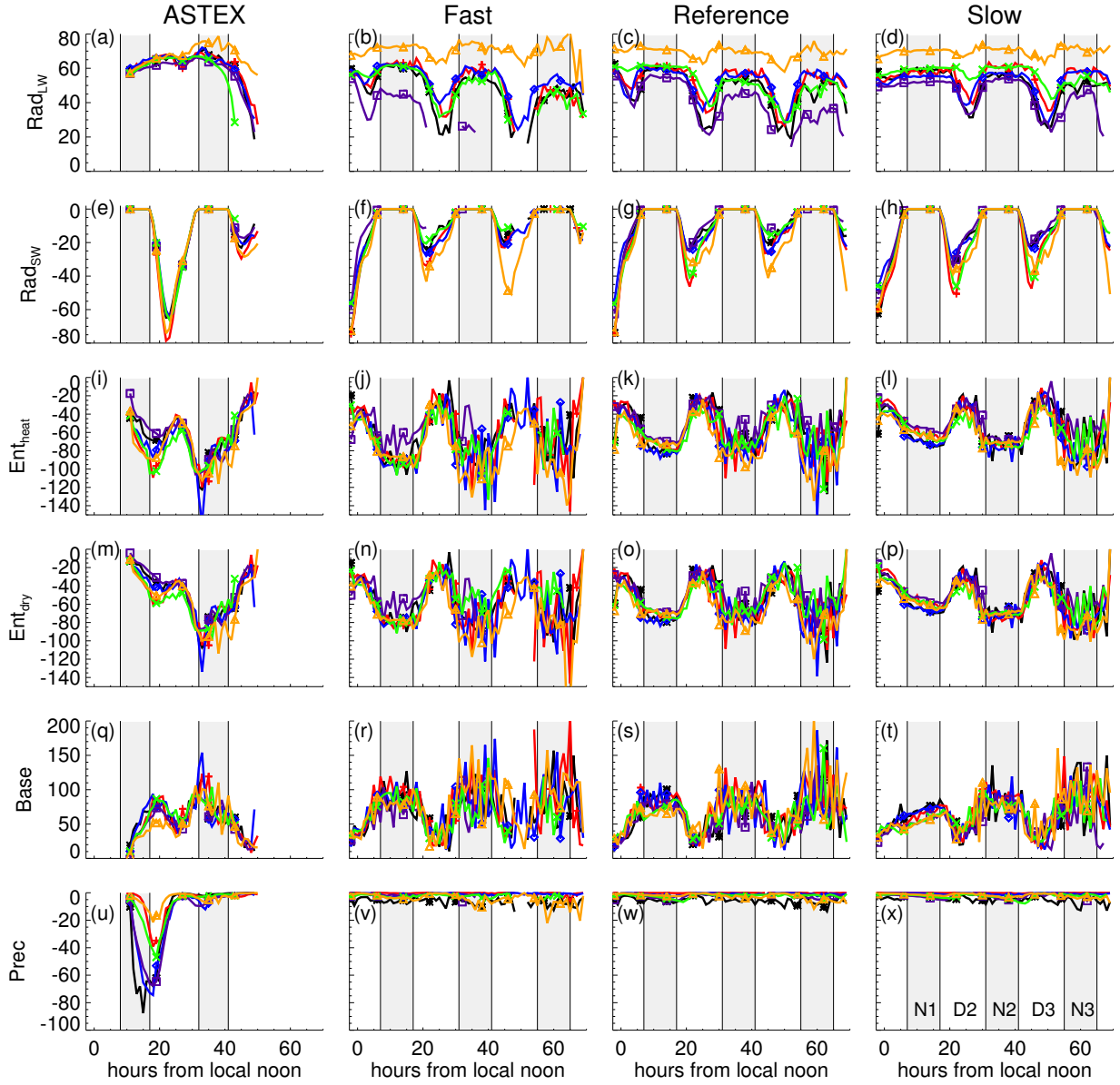


FIG. 12. Time evolution of the dominant terms in the LWP budget, with the variables displayed on the vertical axes denoting LWP tendencies in units of  $\text{g m}^{-2}\text{h}^{-1}$  due to (a-d) longwave radiative cooling ( $\text{Rad}_{\text{LW}}$ ) and (e-h) the absorption of solar radiation in the cloud layer ( $\text{Rad}_{\text{SW}}$ ), entrainment of (i-l) warm ( $\text{Ent}_{\text{heat}}$ ) and (m-p) dry inversion air ( $\text{Ent}_{\text{dry}}$ ), (q-t) cloud base fluxes of heat and moisture ( $\text{Base}$ ), and (u-x) represent the LWP tendency due to drizzle ( $\text{Prec}$ ). The grey shaded bands indicate nighttime periods according to Table 3. The line colors and symbols are as in the legend shown in Fig. 4a.

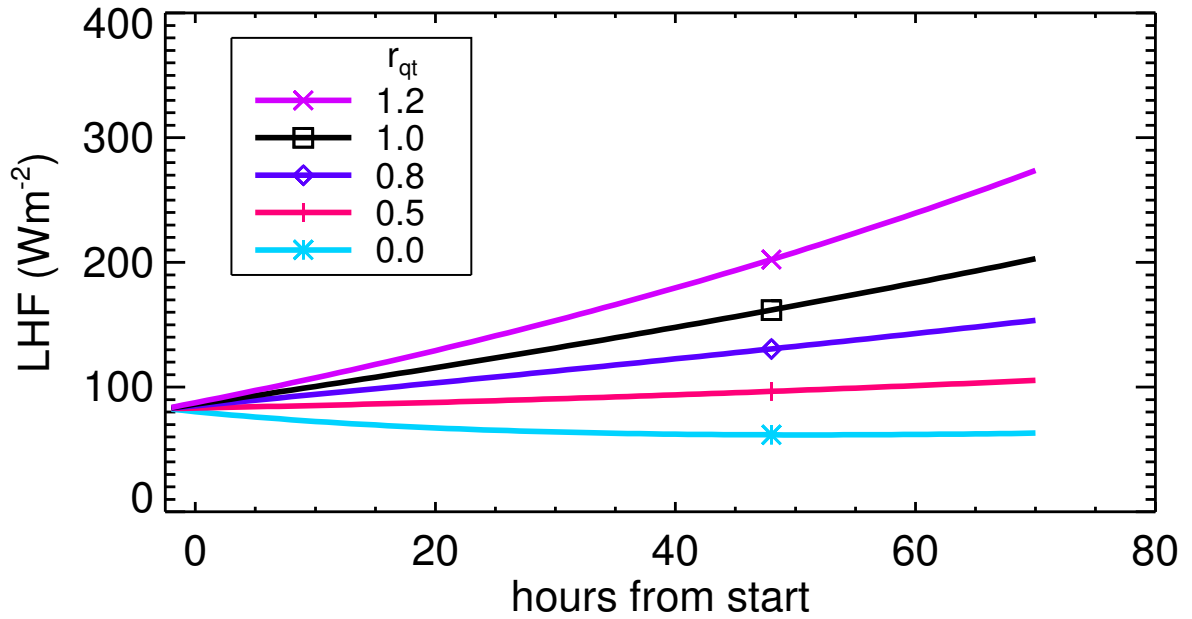


FIG. 13. The latent heat flux as a function of time and for different values of  $r_{qt}$ , which measures the ratio of the total humidity flux at the top of the subcloud layer to its surface value. The line styles are according to the legend.

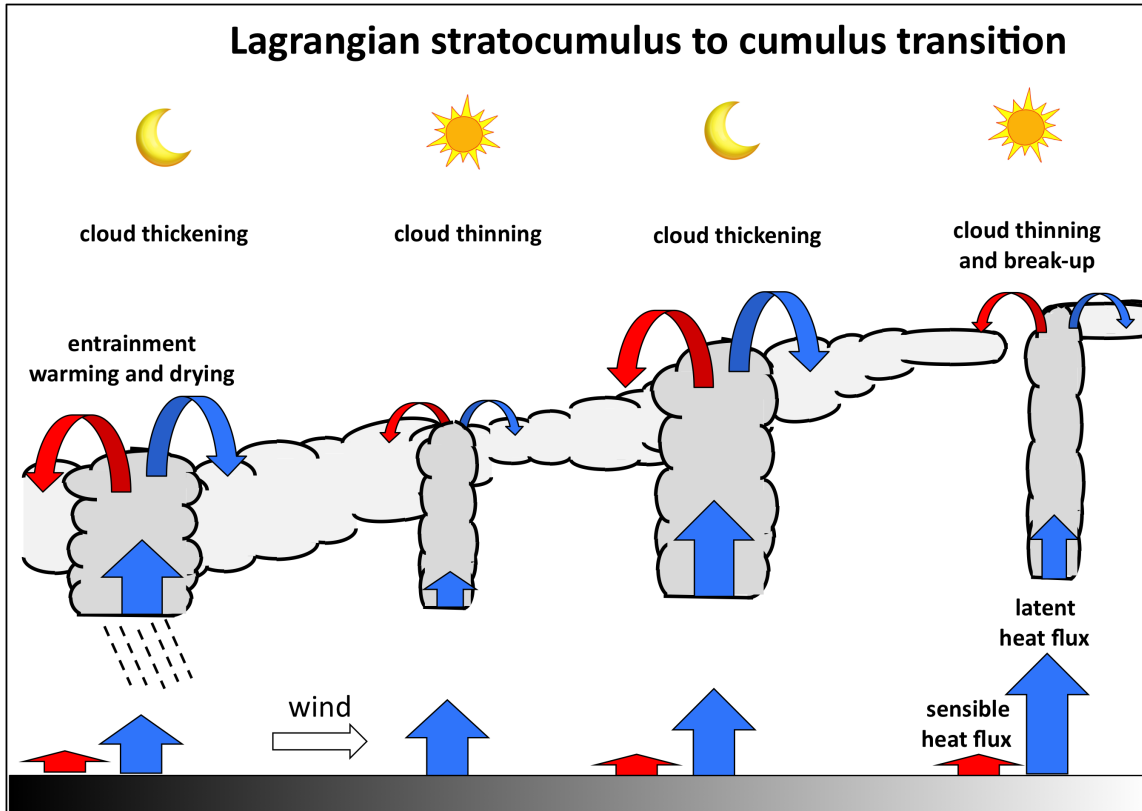


FIG. 14. Schematic showing the gradual break up of a stratocumulus cloud layer during its Lagrangian advection over an increasing SST. The vertical arrows represent the sensible and latent heat fluxes. During the night turbulence in the cloud layer intensifies, causing larger humidity fluxes at cloud base and cloud top.